

An aerial photograph of the Arecibo radio telescope, a large spherical structure with a complex lattice of steel beams, situated in a deep, forested valley. The surrounding landscape is lush green with dense trees and a winding road. The sky is clear and blue.

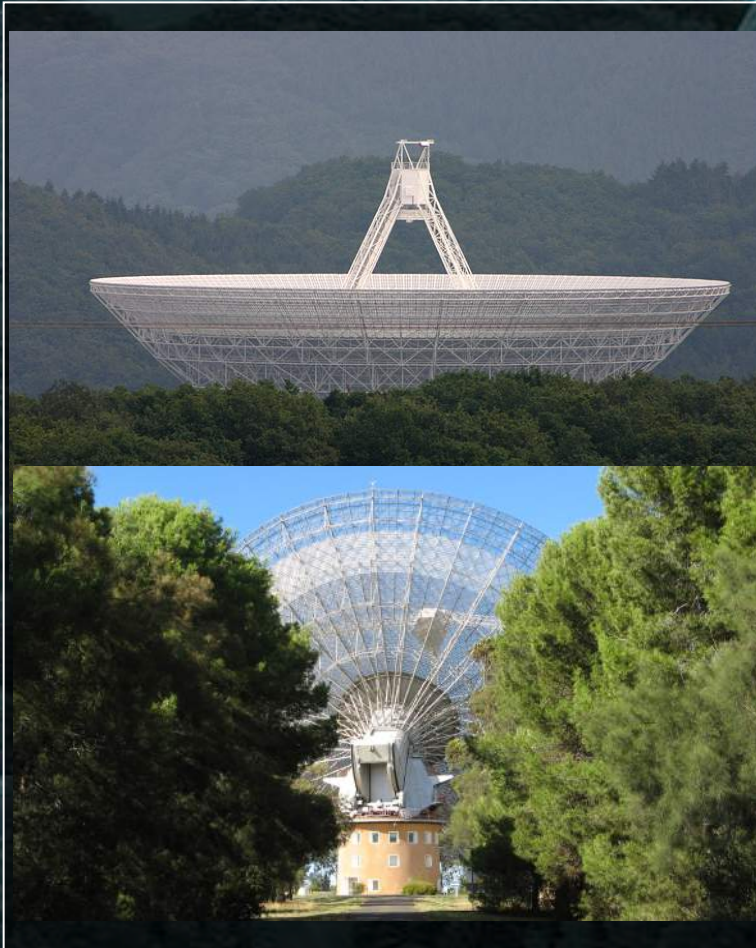
Radio transients in the High Time Resolution Universe radio survey

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On behalf of:
The High Time Resolution Universe consortium

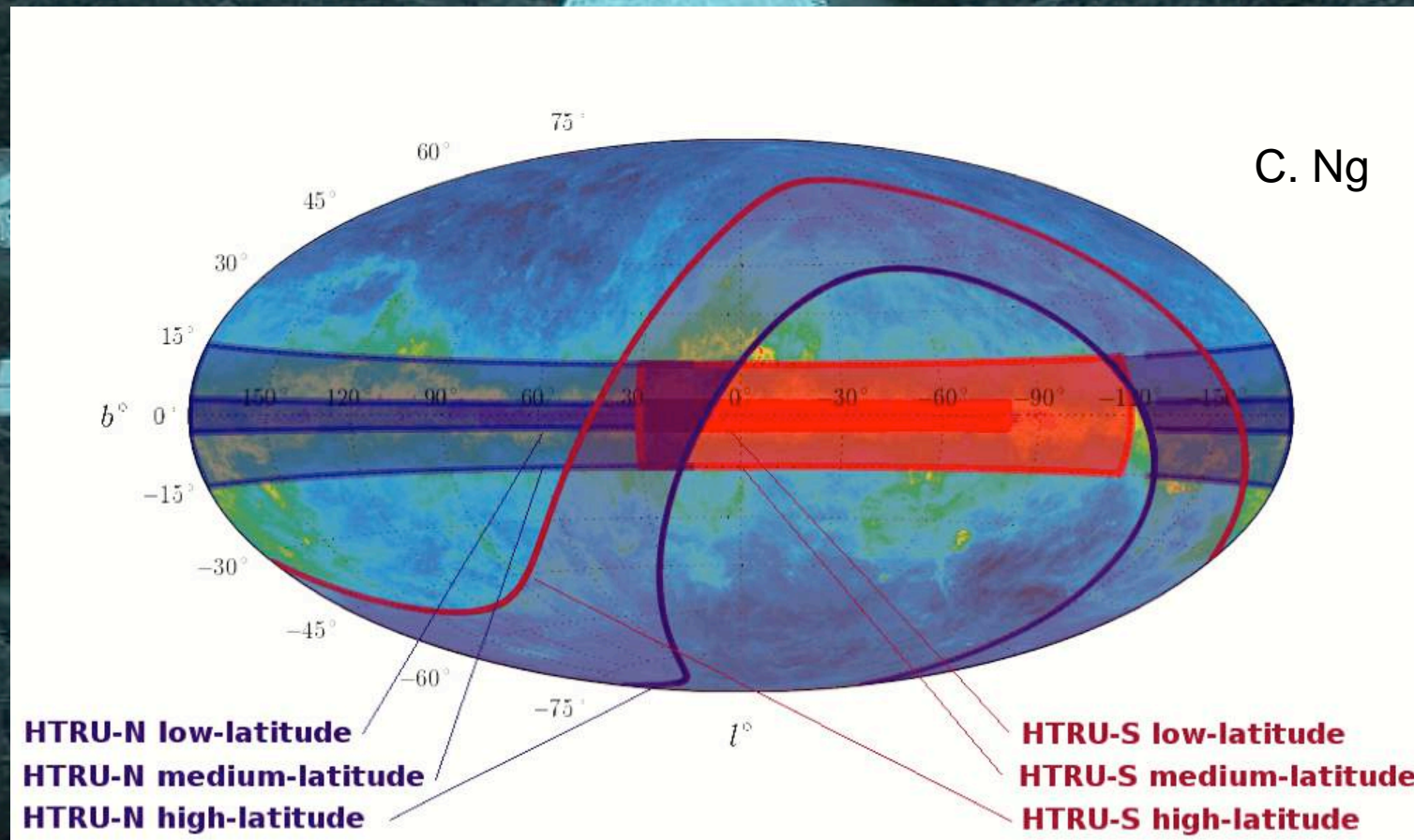
7th Bonn Workshop, 18th May 2015

HTRU comparison with PMPS



	PMPS	HTRU-N	HTRU-S
Start date:	1997 - 2003	Summer 2010	Early 2008
Telescope:	Parkes-64m	Effelsberg-100m	Parkes-64m
Sky coverage:	$\delta < +10^\circ$	$\delta > 0^\circ$	$\delta < +10^\circ$
Integration time:	2100 s	Low-lat: 1500 s Mid-lat: 180 s High-lat: 90 s	Low-lat: 4300 s Mid-lat: 540 s High-lat: 270 s
Receiver:	13-beam 1.35-GHz receiver	7-beam 1.4-GHz receiver	13-beam 1.35-GHz receiver
Backend:	Analogue filterbank (AFB)	Pulsar Fast Fourier Transform Spectrometer (PFFTS)	Berkeley-Parkes-Swinburne Recorder (BPSR)
Bandwidth:	288MHz	300MHz	340MHz
No. of channels:	96	512	1024
Freq resolution:	3MHz	0.58MHz	0.39MHz
Time resolution:	250 μ s	54 μ s	64 μ s
No. sky pointings:	~ 3,200	~ 180,000	~ 43,000
Data sizes:	~ 4 terabytes	~ 5 petabytes	~ 1 petabyte

HTRU search areas



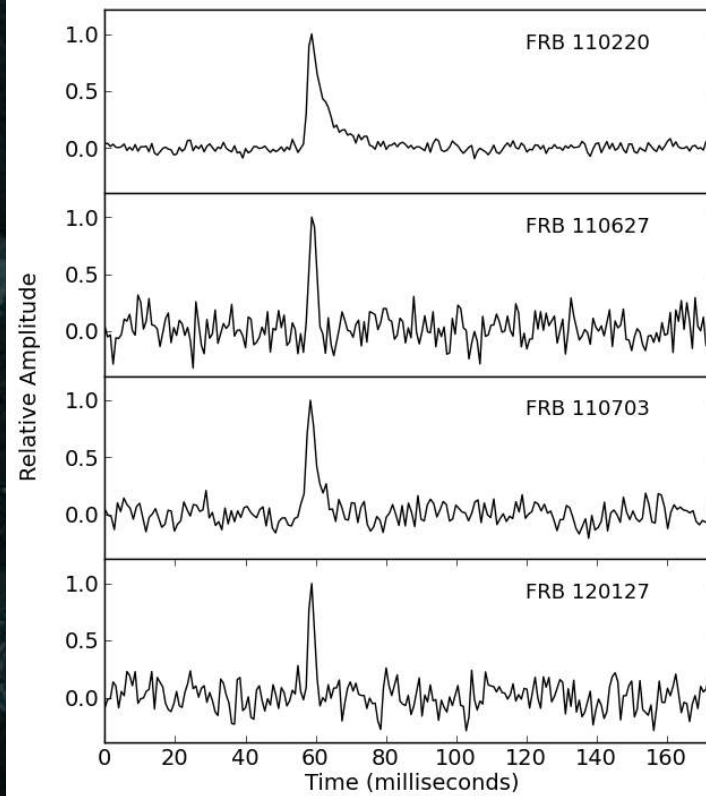
FRB discoveries

Title: Discovery of a fast radio burst population at cosmological distances

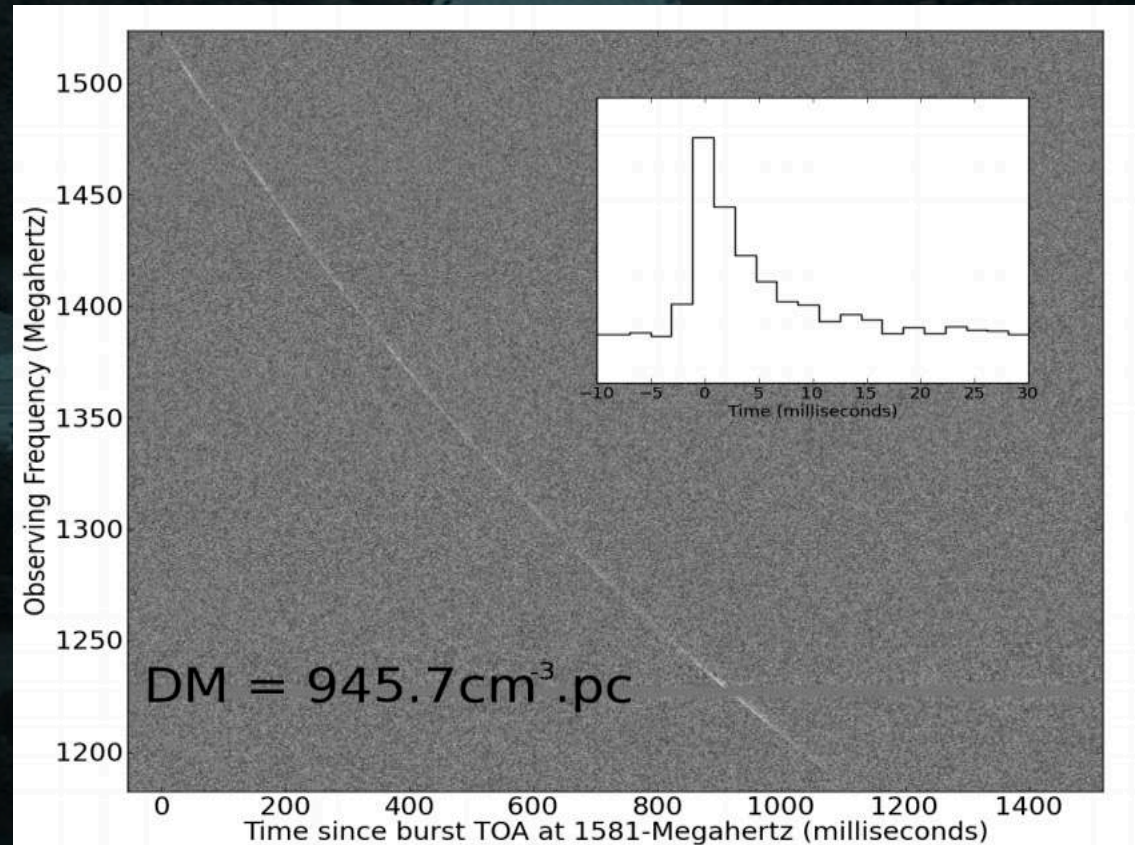
Authors: D. Thornton¹, B. Stappers¹, M. Bailes^{2,3}, B. Barsdell², S. Bates⁴, N. D. R. Bhat^{2,3,5}, M. Burgay⁶, S. Burke-Spolaor⁷, D. Champion⁸, P. Coster², N. D'Amico⁶, A. Jameson², S. Johnston^{2,9}, M. Keith⁹, M. Kramer^{1,8}, L. Levin², S. Milia⁶, C. Ng⁸, A. Possenti⁶, W. van Straten^{2,3}

FRB: Fast Radio Burst, 4 detections

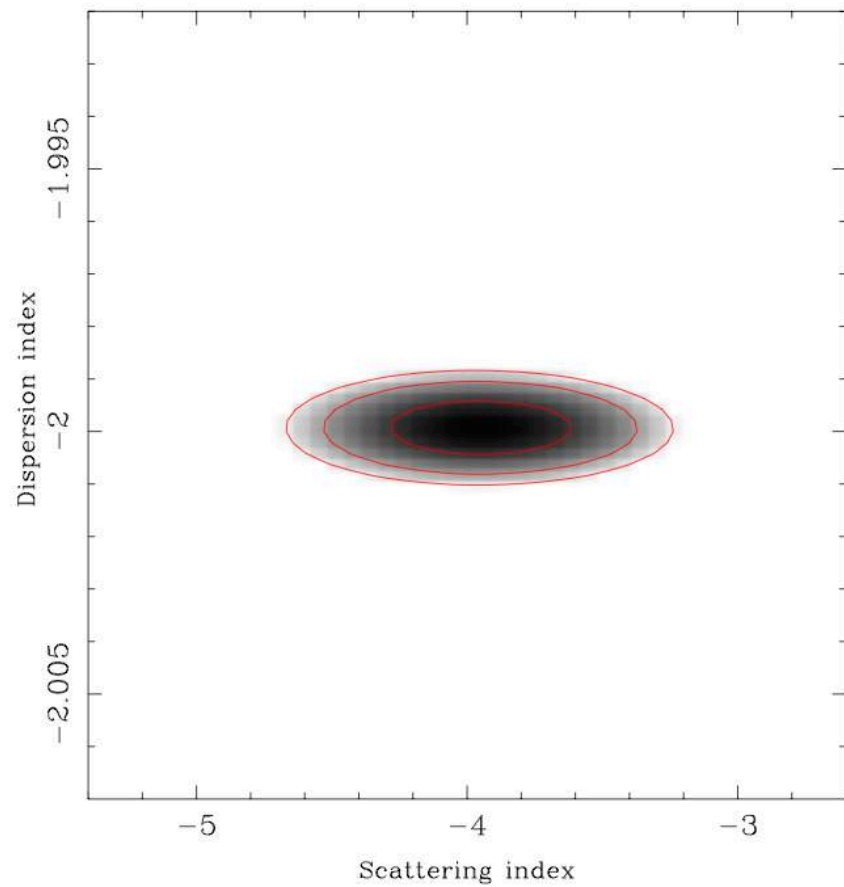
4 FRBs



FRB 110220



DM and Scattering for FRB 110220



Preliminary

5 more...

Name	Position (J2000)	DM (pc cm ⁻³)	DM _{Gal} (pc cm ⁻³)	Width (ms)	Scattering (ms)
FRB090625	03:07:47-29:55:36	897	32	<1.95	<0.98
FRB121002	18:14:47-85:11:53	1629	74	1.3 (5), 2.8 (5), [3.3 (5)]	0.3 (1)
FRB130626	16:27:06-07:27:48	952	67	0.73 (12)	1.09 (11)
FRB130628	09:03:02+03:26:16	469	53	0.39 (5)	0.17 (2)
FRB130729	13:41:21-05:59:43	860	31	<2.93	<1.5

- All data (including Thornton) processed using Heimdall
- All previously known FRBs redetected
- No FRBs missed in old data

$$S/N \geq 10$$

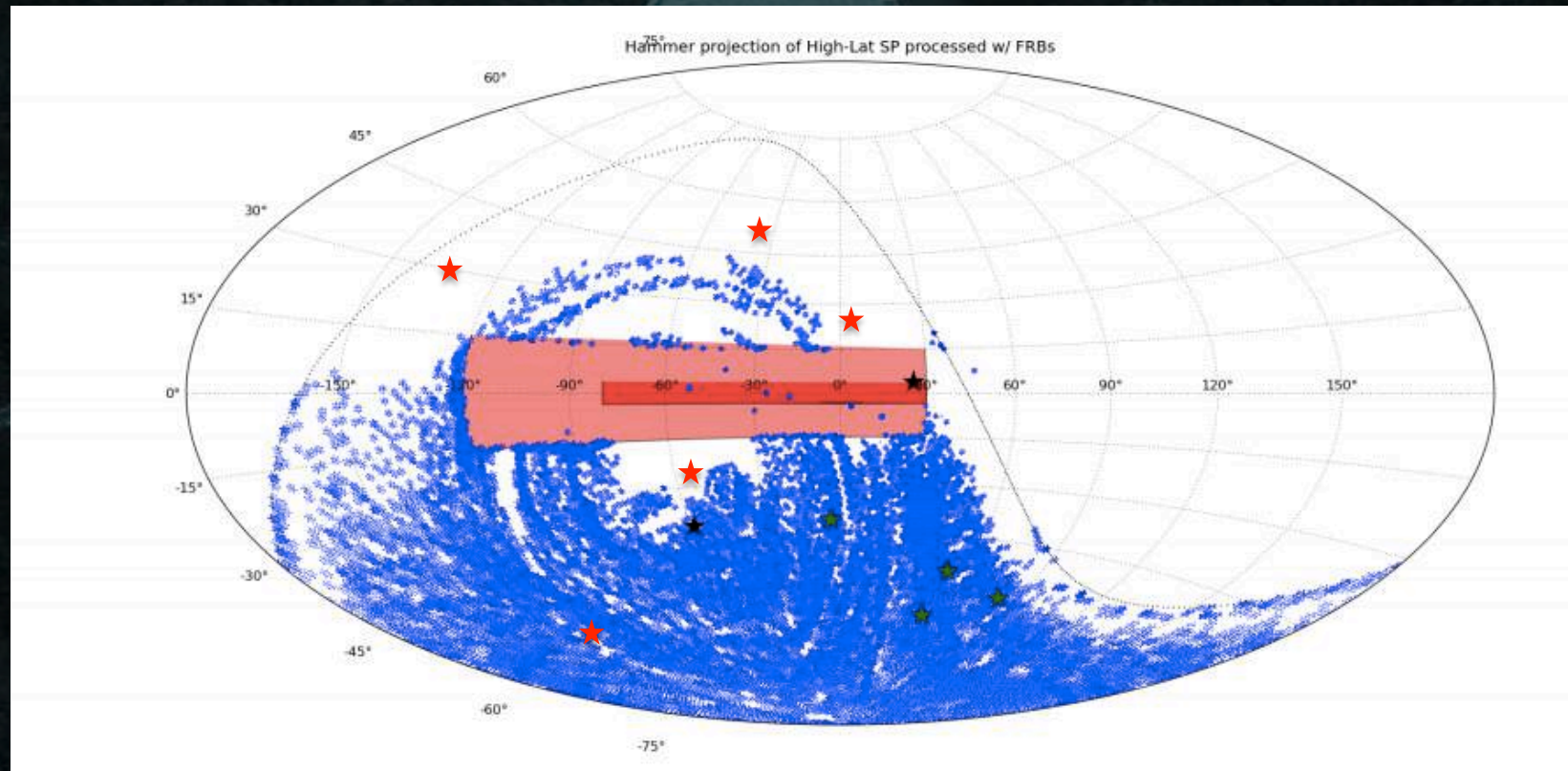
$$\Delta t \leq 2^8 \times 64 \mu\text{s} = 16.3 \text{ ms}$$

$$DM/DM_{\text{Galaxy}} > 0.9$$

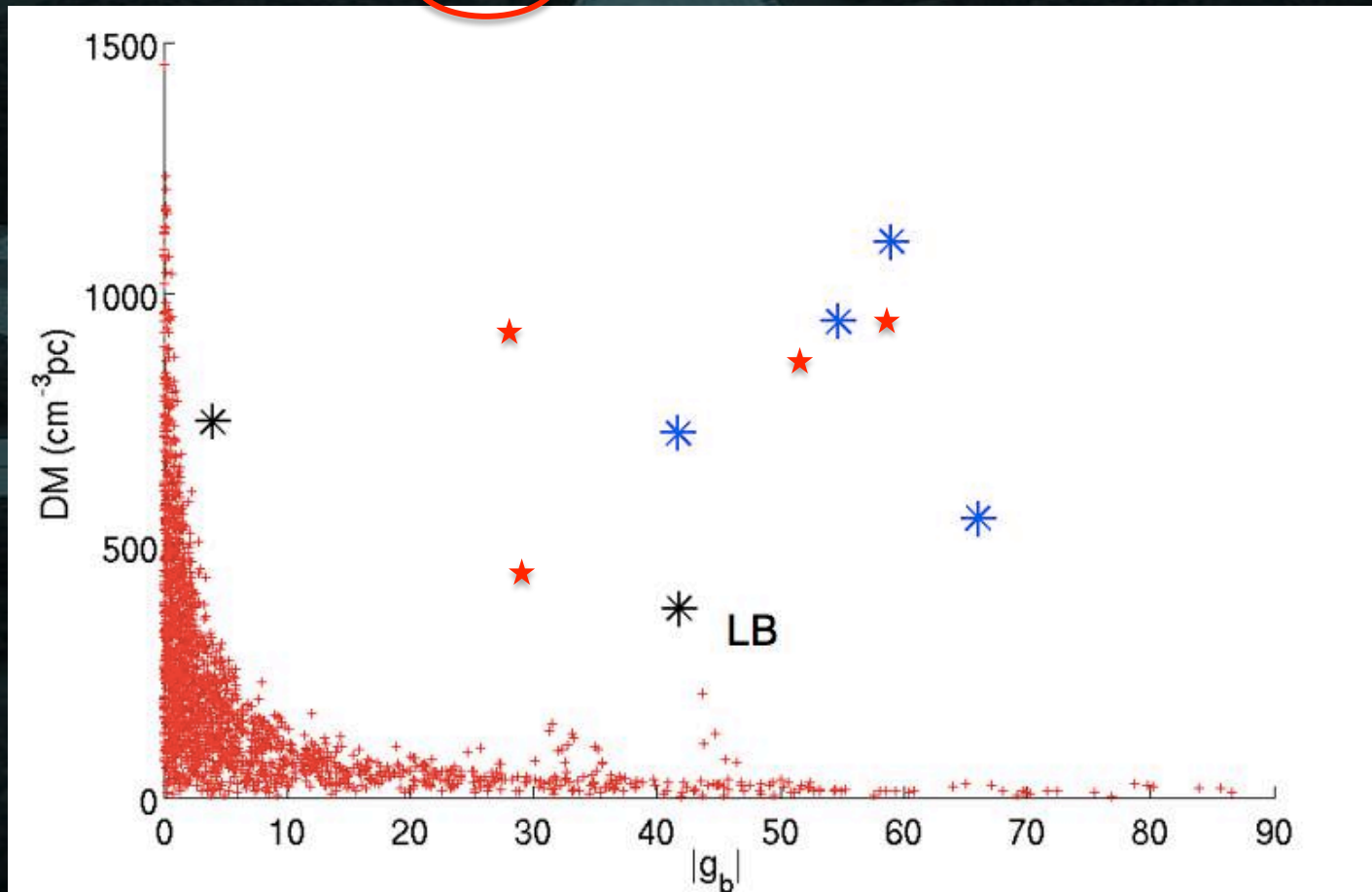
$$N_{\text{beams}} \leq 4$$

Champion et al in prep.

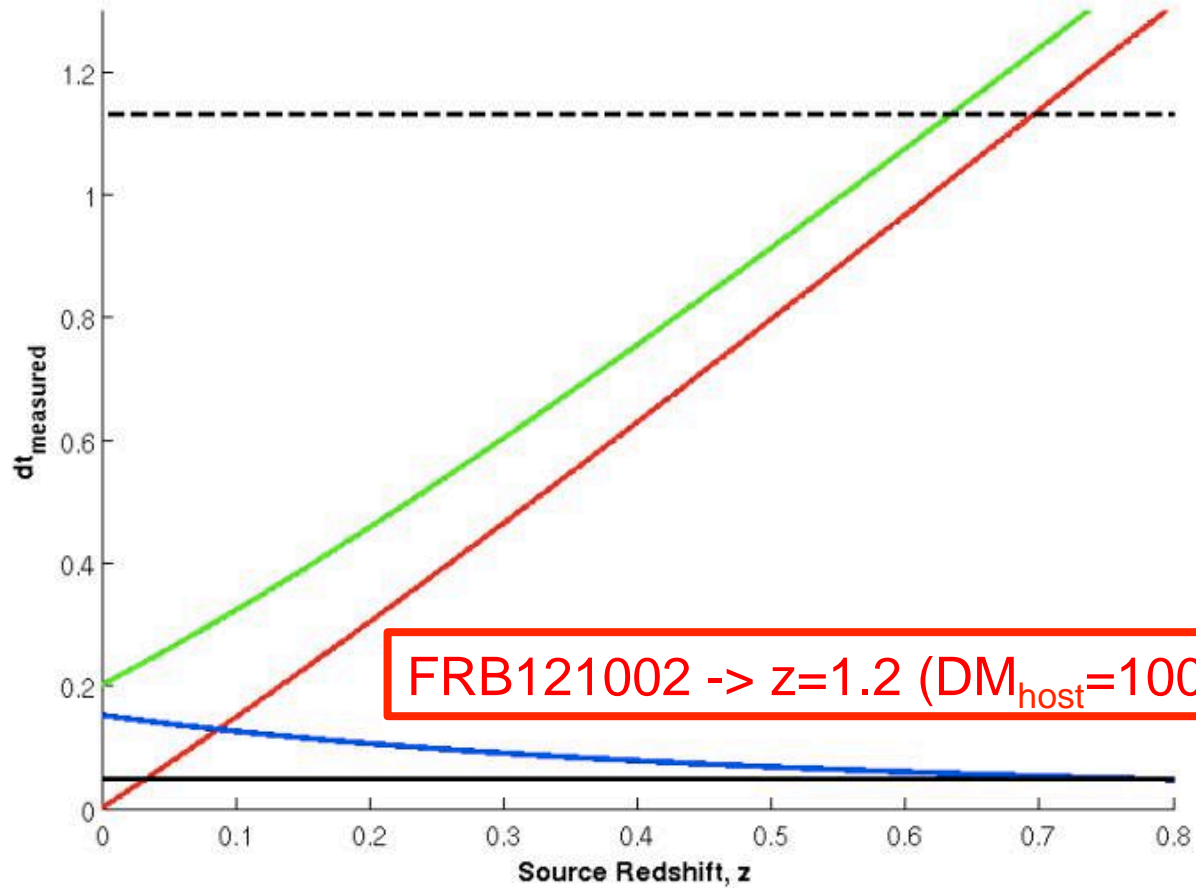
Locations of the bursts



DMs of the bursts



DM and redshift



Preliminary

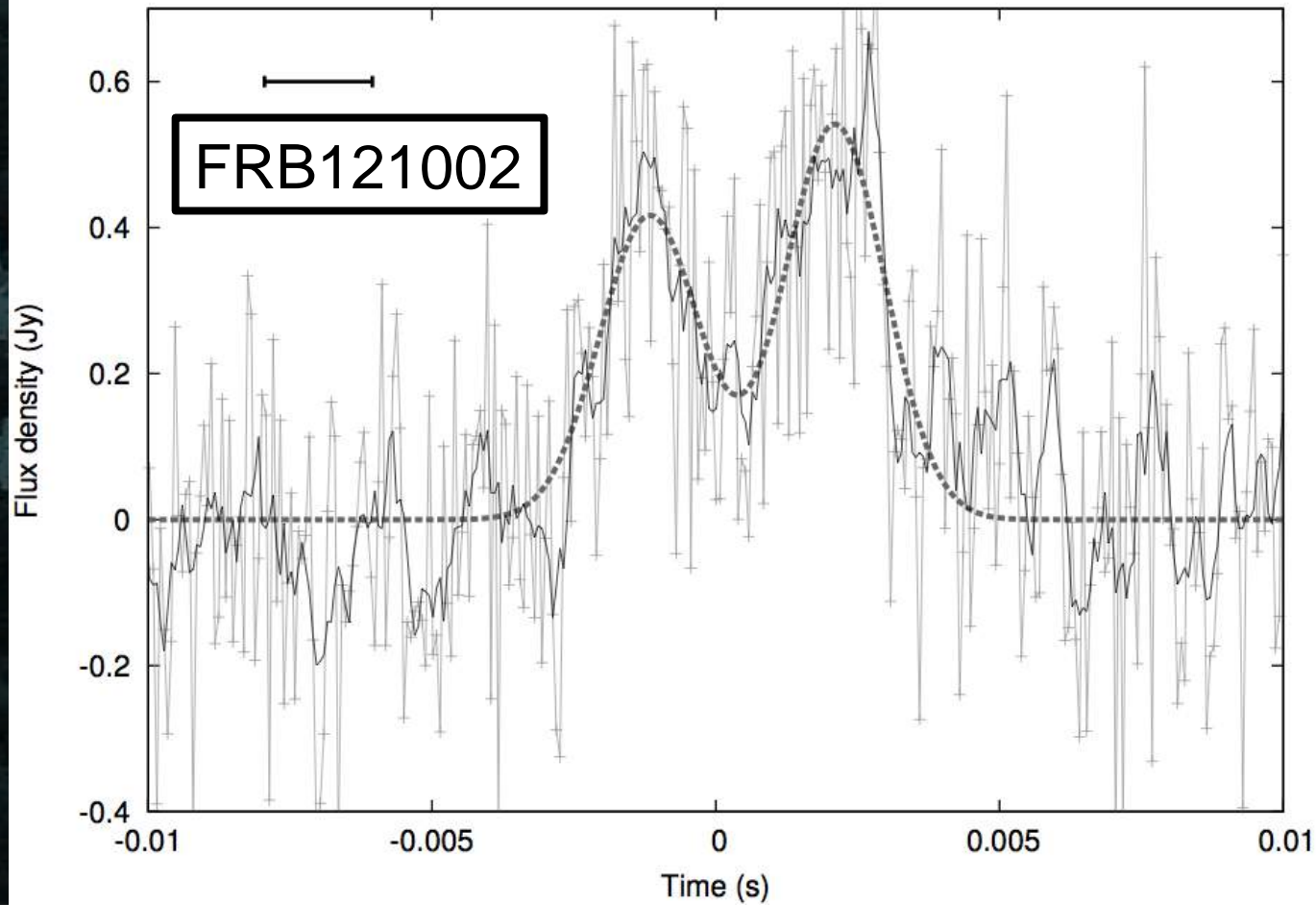
Rate of FRBs

- 100% of HTRU-S high-lat checked:
33,500 pointings x 13 beams x 270 sec = 1549 deg²hrs

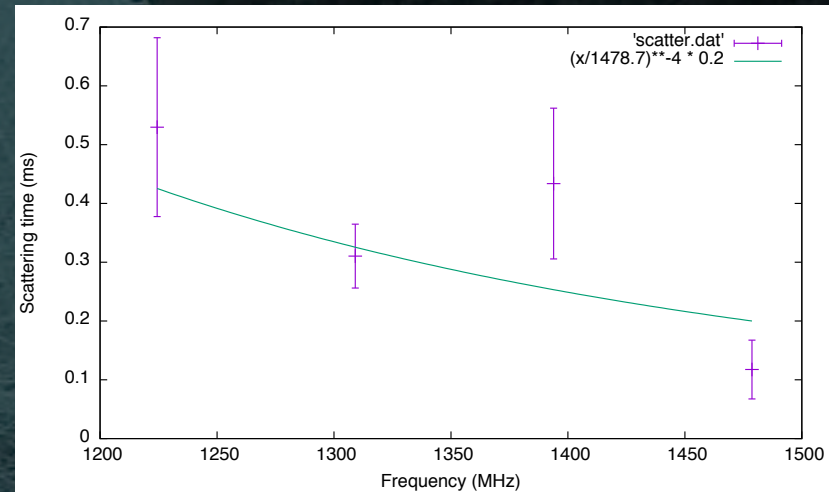
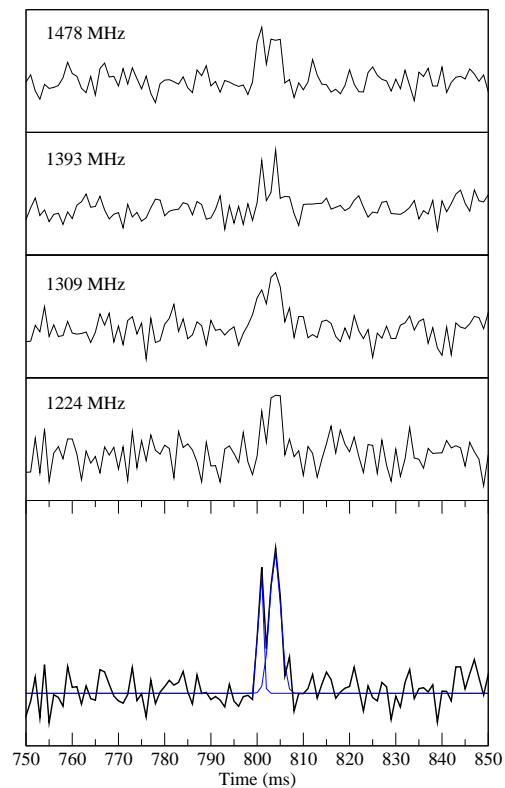
$$9 \text{ FRBs} \times \frac{24 \text{ hrs/day} \times 41253 \text{ deg}^2/\text{sky}}{1549 \text{ deg}^2 \text{ hrs}}$$

- $R_{\text{FRB}}(>0.6\text{Jy ms}) \sim 5.7 \times 10^3 \text{ }^{+4.2}_{-2.7} \text{ sky}^{-1} \text{ day}^{-1}$
- At lower end but consistent with Thornton et al 2013
- Consistent with Spitler et al. 2014
- 97% inconsistent with MedLat rate (Petroff et al 2014)

Double component FRB



Double component FRB

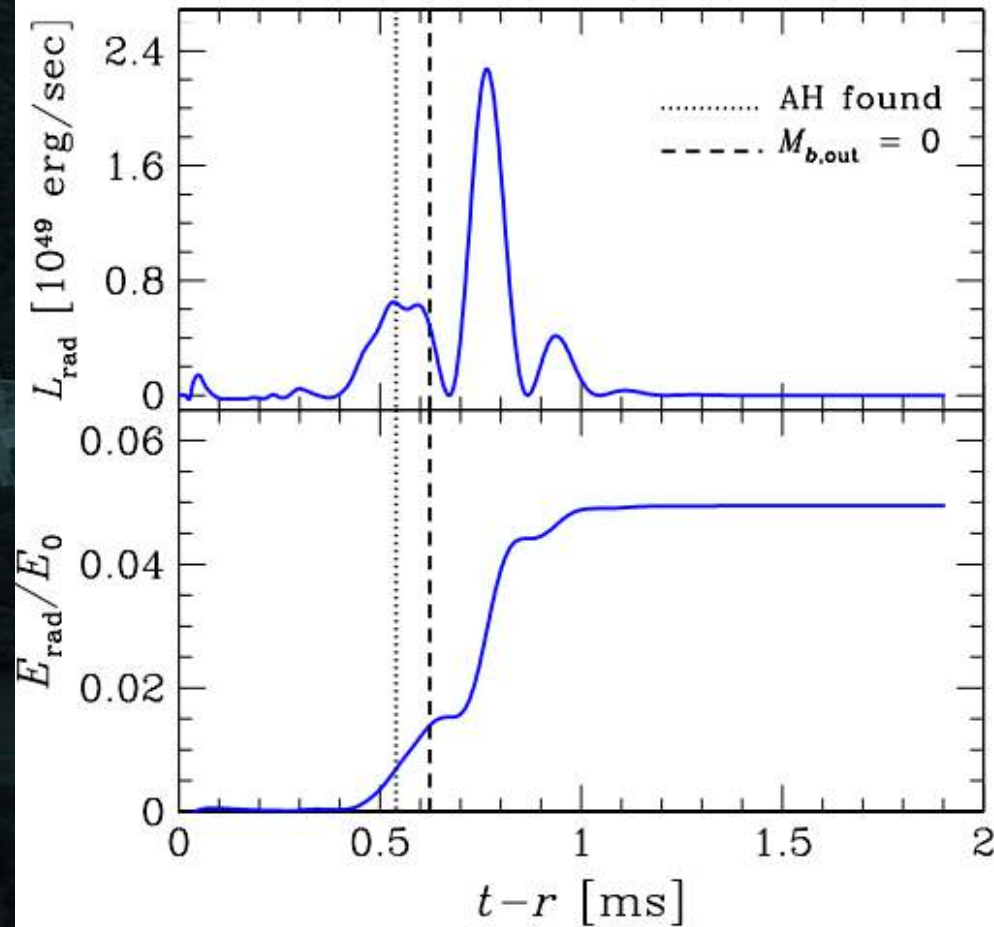


Subbands fitted independently
Highest subband used as template
Fitted for scattering and amplitude
Consistent with global fit

Models

- Evaporation of BHs (Rees 1977)
 - WD merger (Kashiyama et al 2013)
 - NS merger (Hansen & Lyutikov 2001)
 - Cosmic string collisions (Cai et al. 2012)
-
- SGR giant flares (e.g. Thornton et al. 2013)
 - SGR hyper flares (Popov & Postnov 2007)
 - Super giant pulses (Cordes & Wasserman 2015)
 - Collapse of a supra massive BH (Falcke & Rezzolla 2014)

Supra massive BH collapse



Conclusions

- The HTRU surveys are providing a snapshot of the radio sky for pulsars and transients
- We have confirmed a new population of extragalactic radio transients
- The rate is lower than but consistent with the Thornton et al. rate
- First evidence of structure in FRB
- Challenges some origin theories
- NS sources and supra-massive BH model explain this more easily