# Uncertainties in pre-supernovae stellar models

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Mesa IV: Paxton+,Farmer 2017

> SNII LC PPISN/PISN + more

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### Final fate of a star



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Pejcha & Thompson 2015 Also: Ugliano+ 2012, O'Connor & Ott 2013, Müller+ 2016, Sukhbold+ 2016, 2017, Ertl+ 2016

# Compactness



\*Not to scale

## Compactness



\*Not to scale

### Compactness



Compactness still isn't the complete picture



# What sets the compactness?



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#### Fields+,Farmer 2017



Uncertainty in a rate varies with temperature





~20% variations just from uncertainties in the things we can measure in a lab

Also: West+2013, Cyburt+2016, Fields+ 2016, Rauscher+ 2016, Nishimura+ 2017, Harris+ 2017



Key rates: 12C→160, 14N→15O, 15N→16O, 3α

Also: West+2013, Cyburt+2016, Fields+ 2016, Rauscher+ 2016, Nishimura+ 2017, Harris+ 2017



(T, Rho, L, R, X, v, omega)

#### Farmer+ 2016



#### Farmer+ 2016



#### Farmer+ 2016



Stellar lifetime depends on spatial resolution

#### Farmer+ 2016



Time

















- Many parts of stellar evolution modelling have uncertainties in them
- Small changes can lead to "branches" in the evolution
- Begun to quantify their effect
  - Nuclear reactions ~20%
  - Spatial Resolution ~50%
  - Correlations between terms?

# Final state is determined by more than a star's initial mass & metalicity