



# MILLISECOND PULSAR POPULATION

***Amruta Jaodand***

11th Bonn Neutron Star Workshop on  
Formation and Evolution Of Neutron Stars

11/12/2017

# Outline

**1. Millisecond Pulsars**

**2. Current Population**

**3. Applications**

**4. Future Prospects**

**4.1. Potential Science Questions**

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# Millisecond Pulsars

Introduction



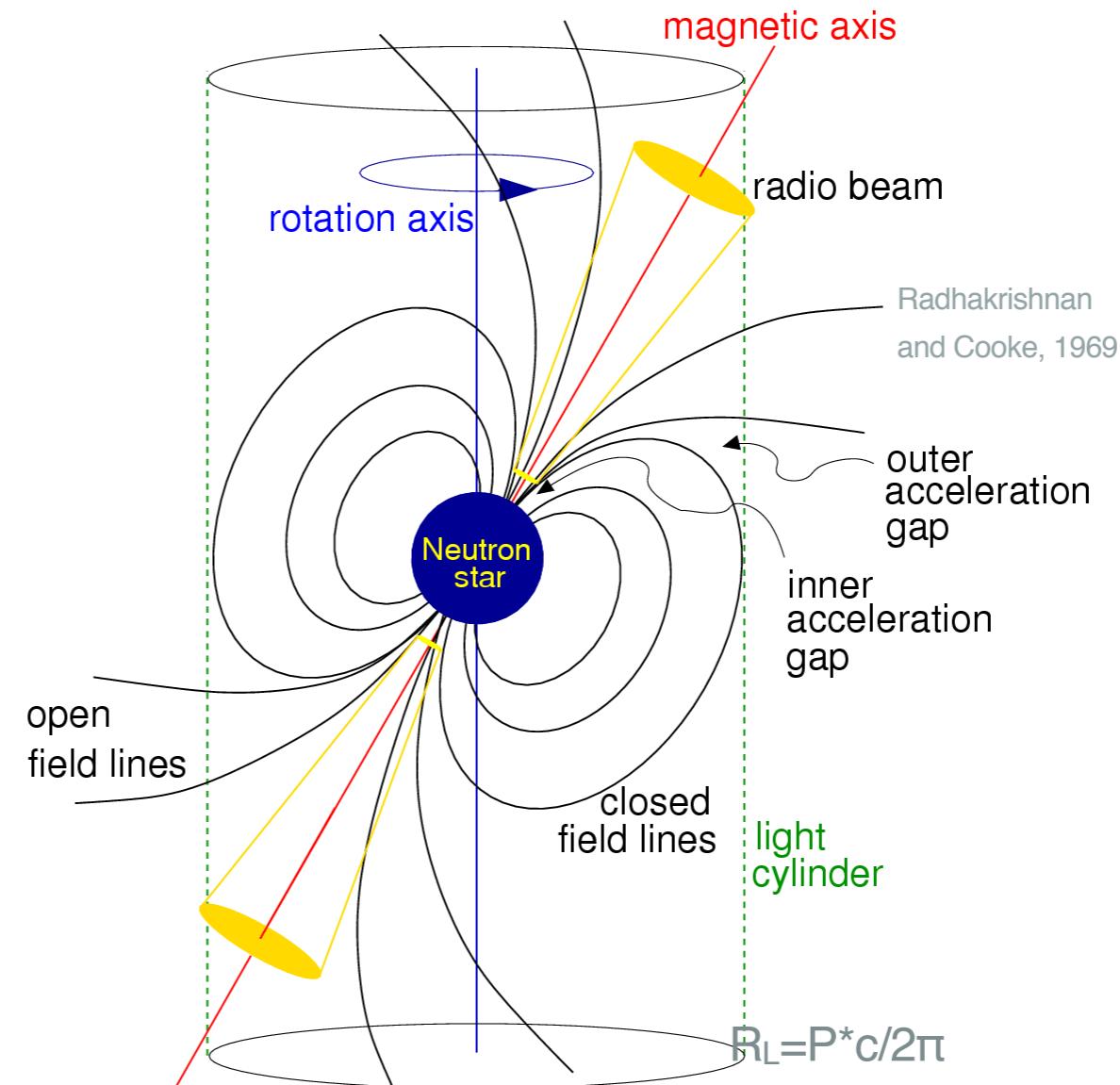
**DISCOVERY**

1982, D. Backer, Millisecond Pulsar, PSR J1937+21

$P < 10\text{ms}$  in this talk.

## DISCOVERY

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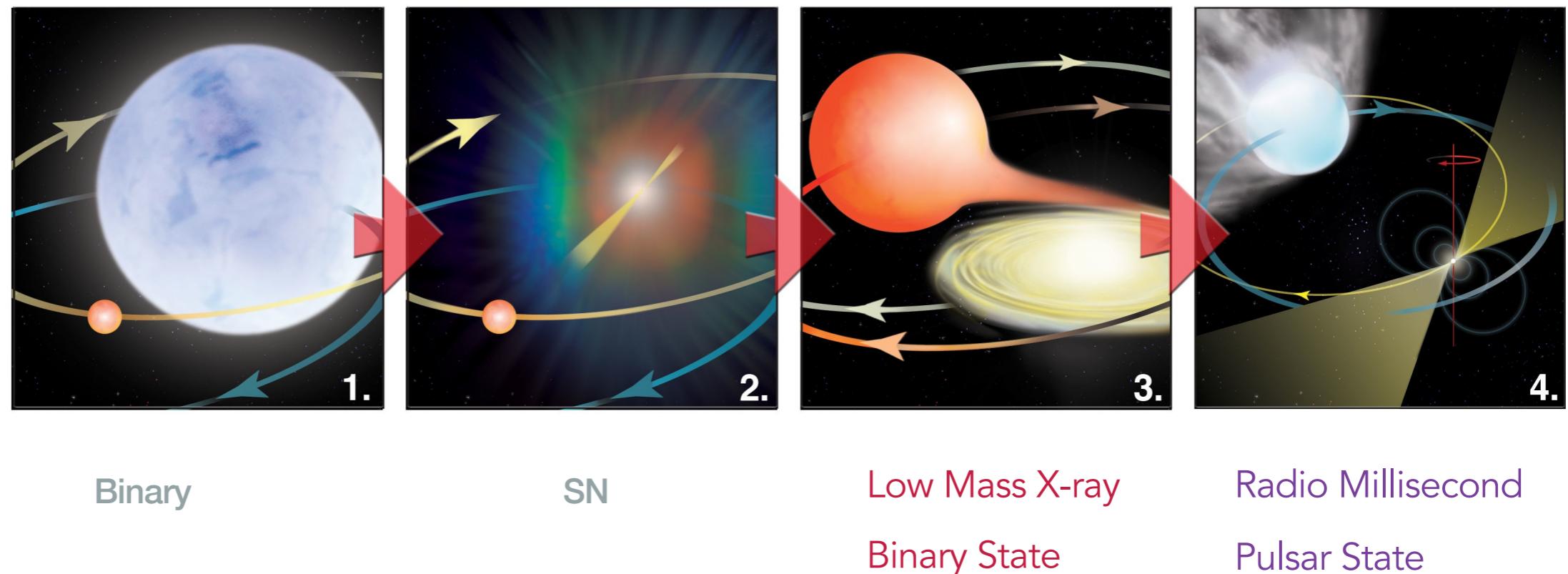
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# PULSAR RECYCLING

Backer, Kulkarni et al. 1982 - Short period, small period derivative

Alpar, Cheng, Ruderman & Shaham, A new class of radio pulsars, Nature, 300, 728, 1982  
Radhakrishnan & Srinivasan 1982

Bill Saxton; NRAO/AUI/NSF

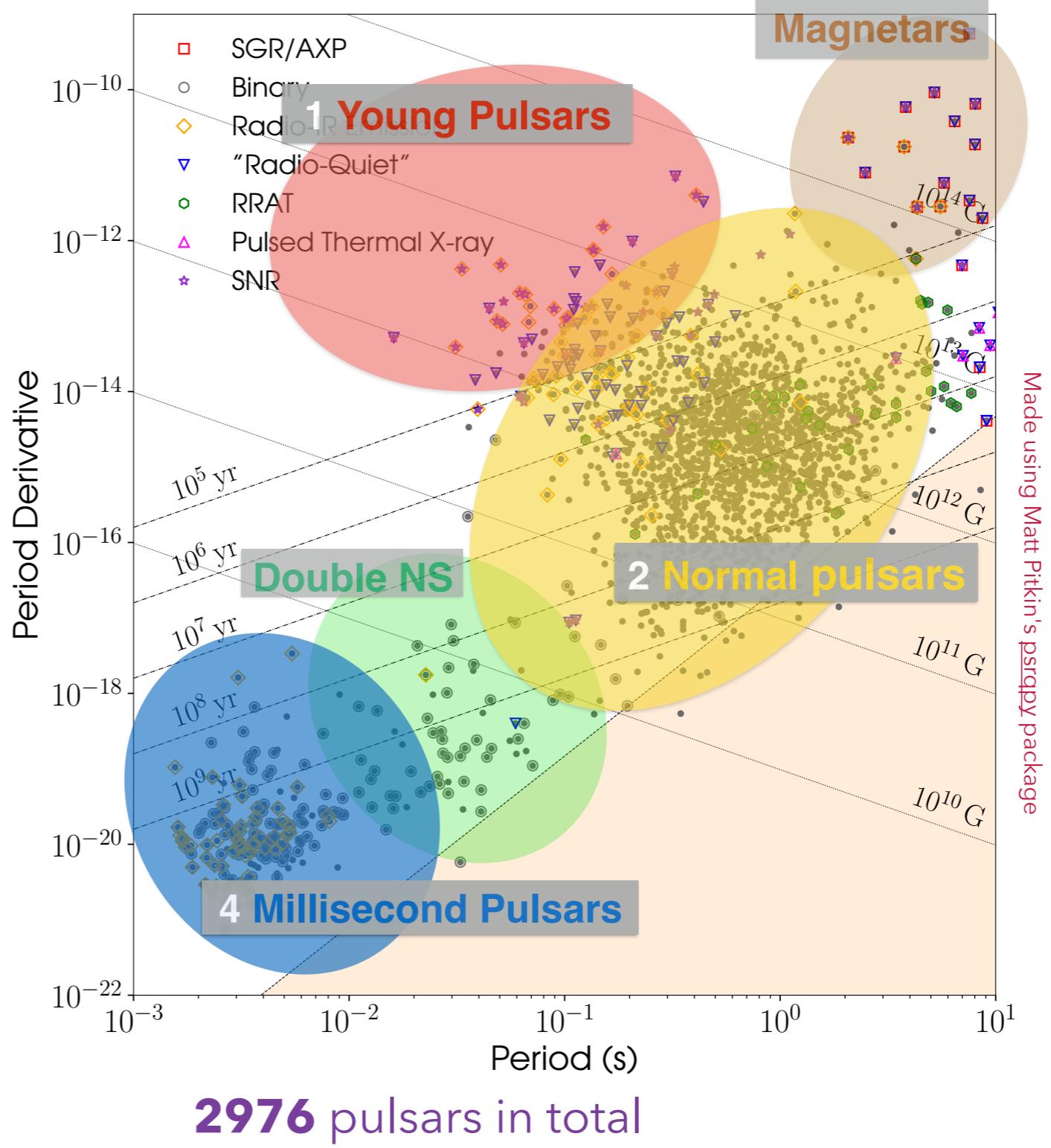


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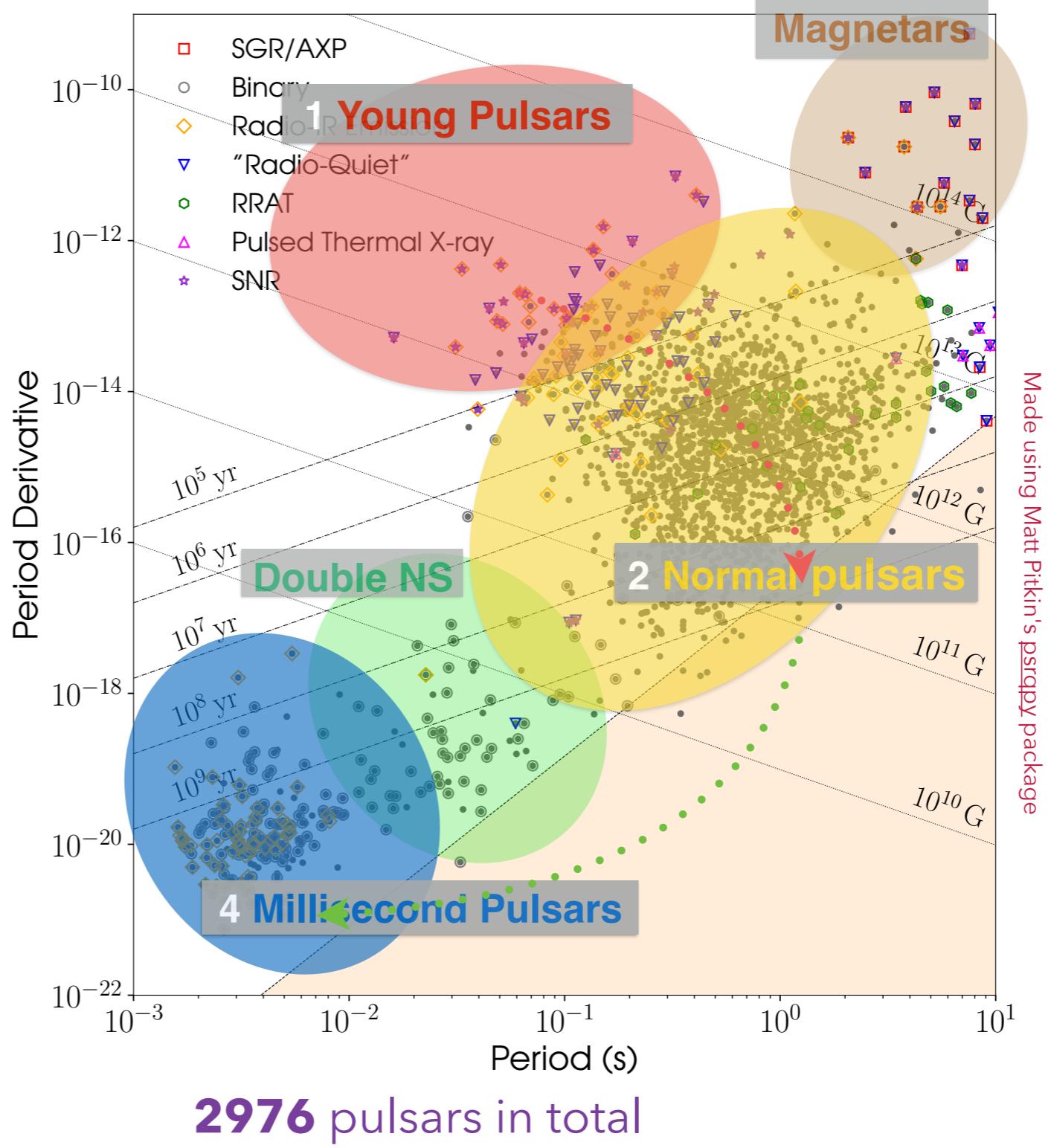
# Current Population

Diverse Manifestations

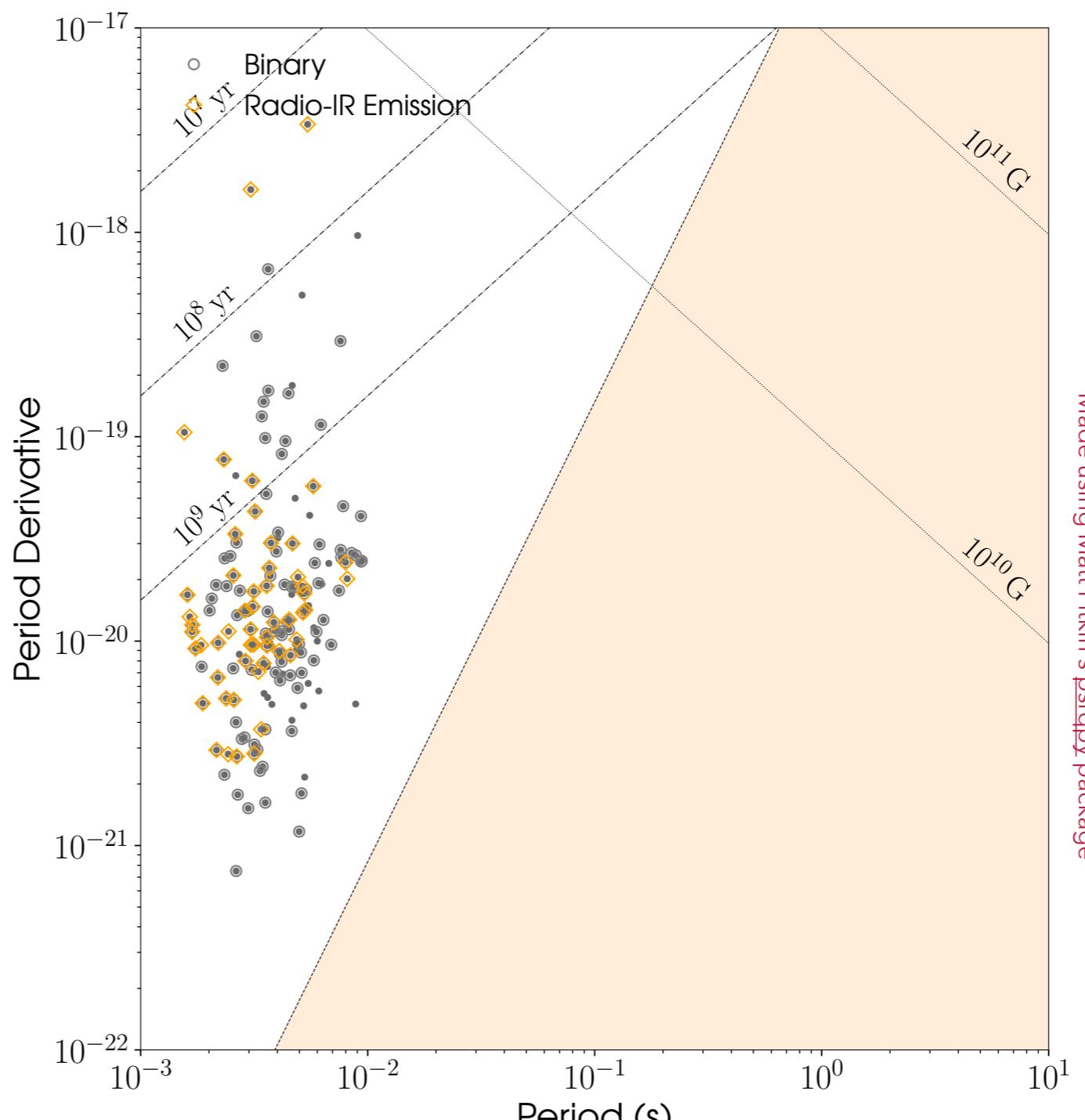
## CURRENT STATUS



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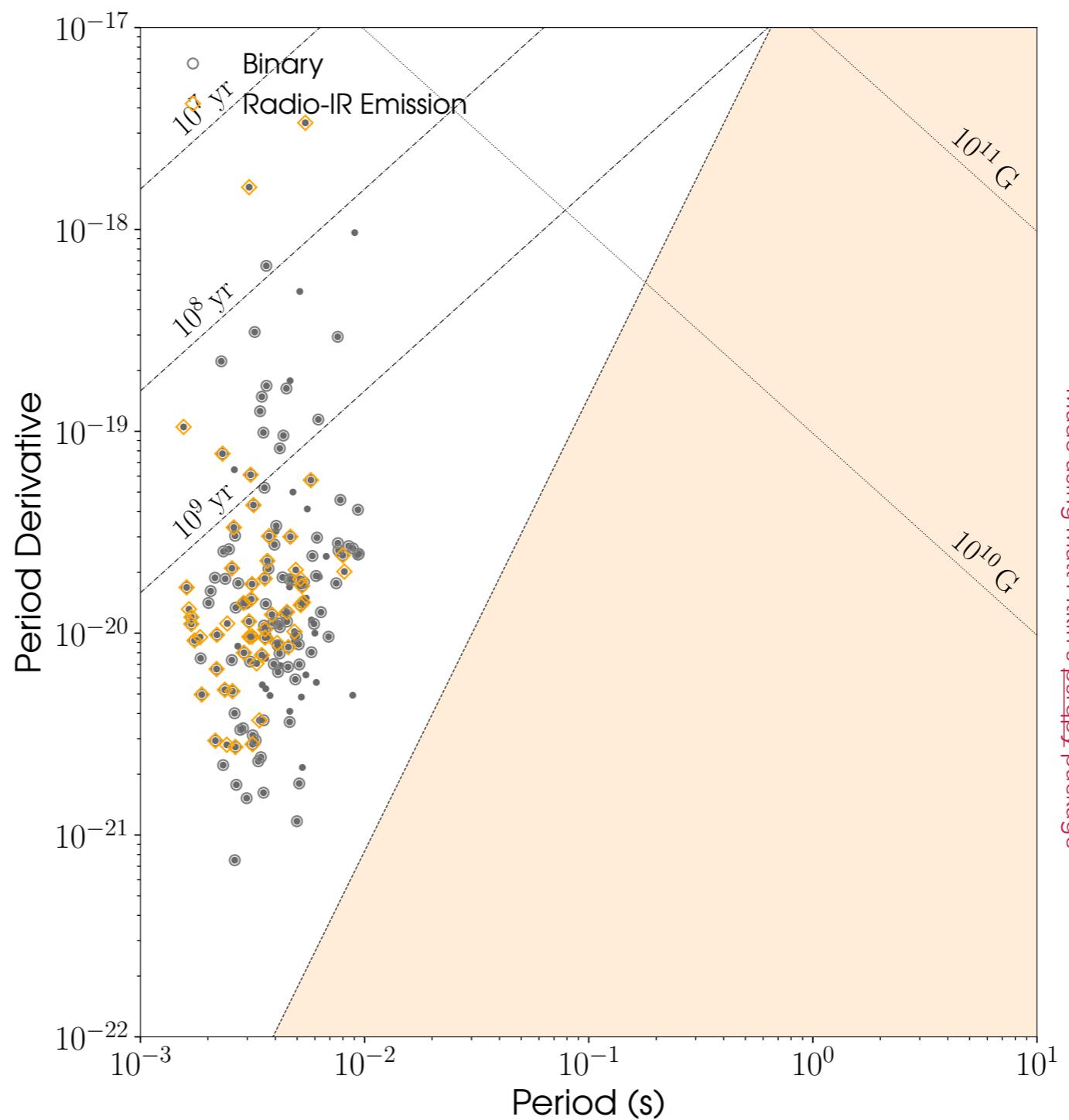
## CURRENT STATUS



**2976** pulsars in total

Made using Matt Pitkin's `psrappy` package

## CURRENT STATUS



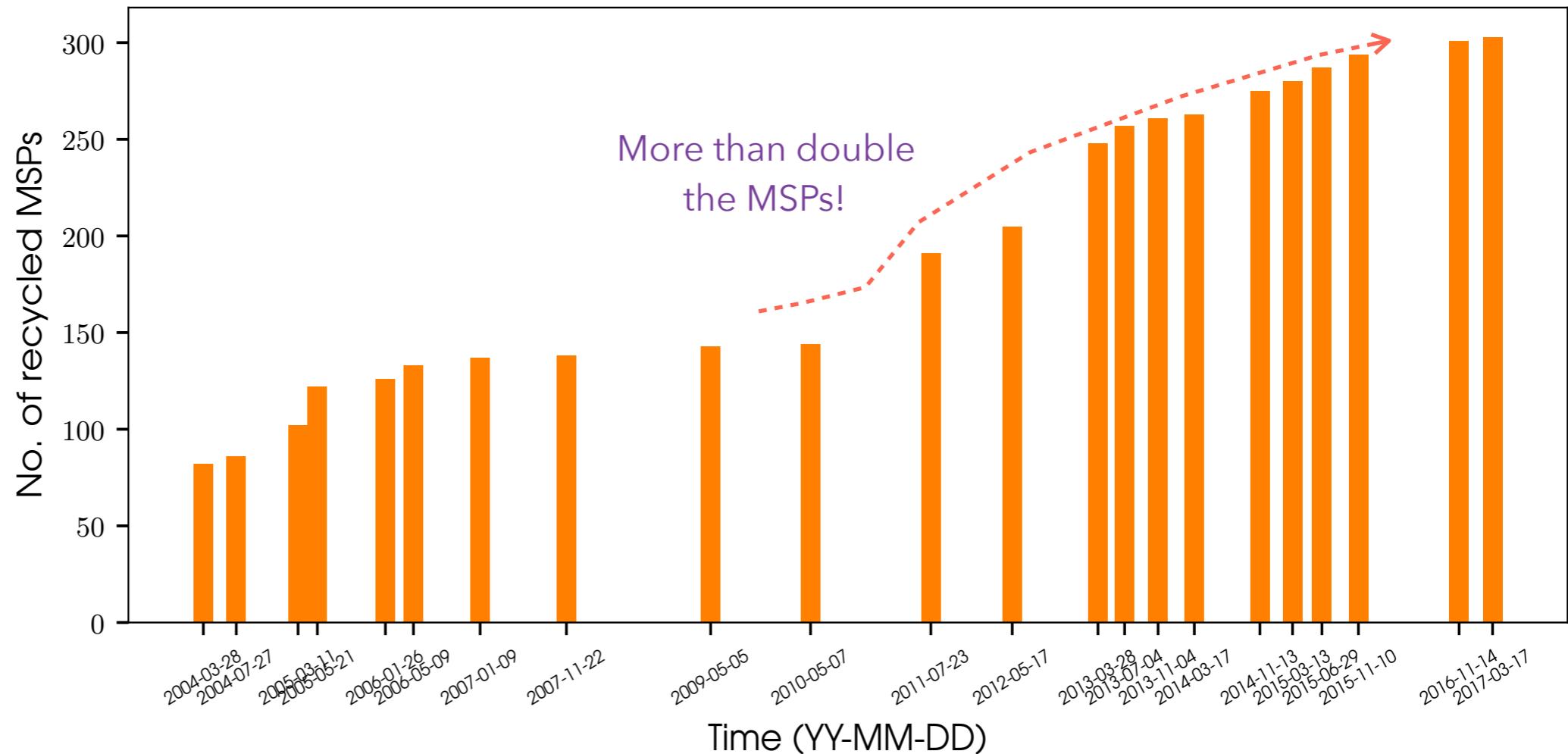
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**308** millisecond pulsars ( $F_0 > 100\text{Hz}$ )

**18** millisecond pulsars ( $F_0 > 500\text{Hz}$ )

**2** millisecond pulsars ( $F_0 > 700\text{Hz}$ )

**PSR BOOM**



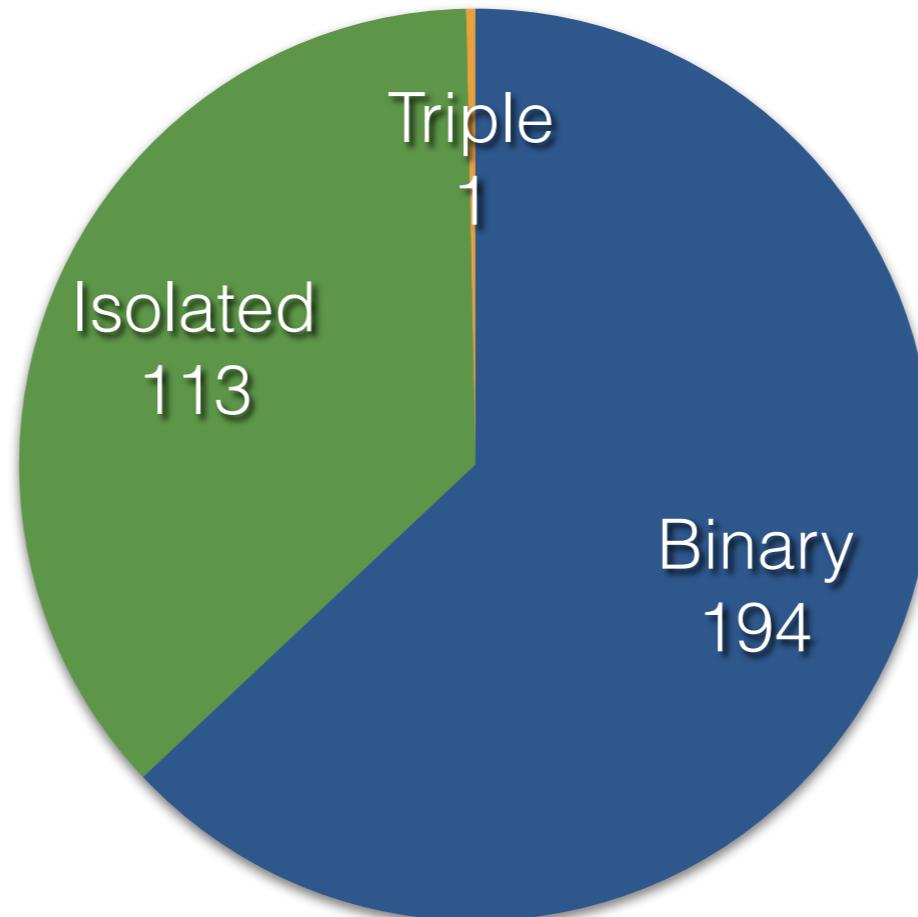
**Sky surveys:** Parkes Multi-Beam led the way!  
GBT-drift scan and North Celestial Cap surveys;  
Arecibo - 327 MHz drift-scan and ALFA survey;  
HTRU- Northern and Southern sky surveys

### *Fermi Treasure Map*

New telescopes, new backends, new algorithms - acceleration searches



## Diversity in MSPs ( $P < 10\text{ms}$ )



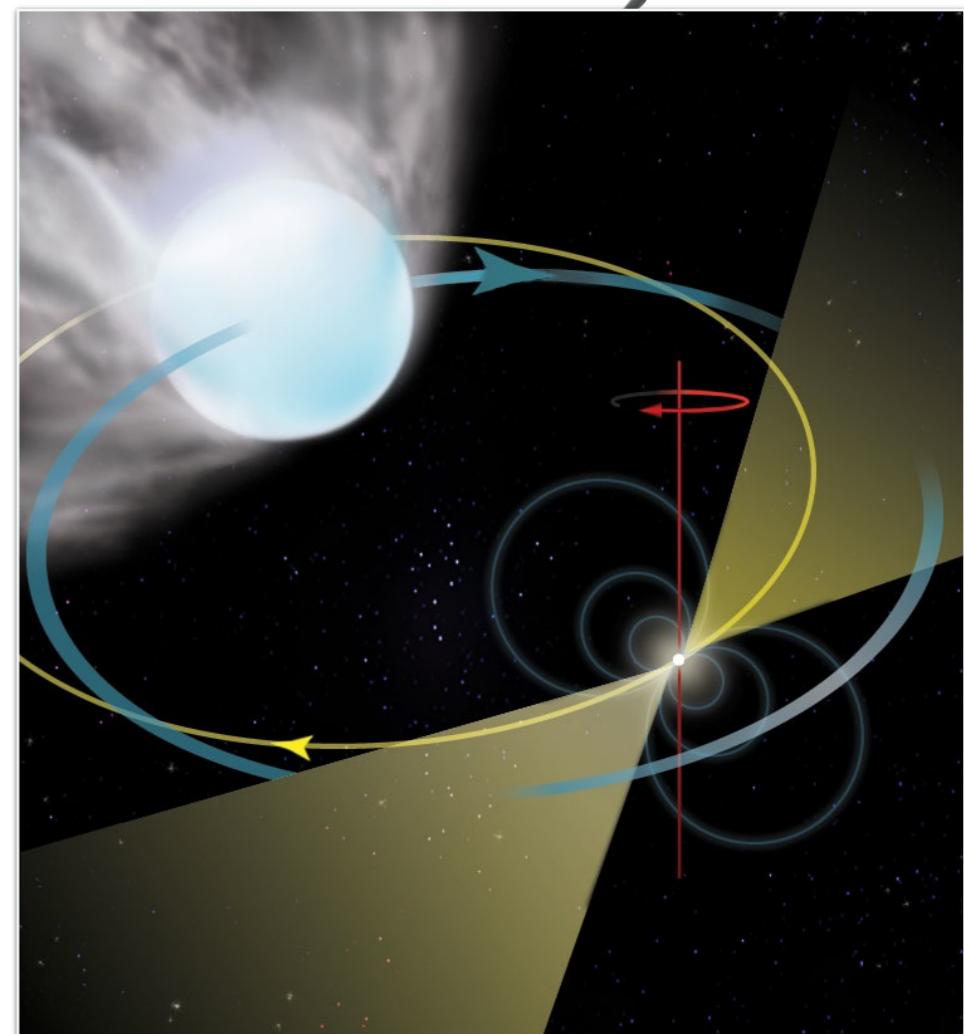
**120 are in Globular Clusters, 116 Radio sources,  
73 Gamma-ray sources, 51 X-ray sources**

MSP  
BINARIES

# Black Widows

“Black Widow” pulsar,  
B1957+20, Fruchter et al.(1988)

Companion Mass,  $<0.01\text{-}0.1 M_{\text{sun}}$   
Orbital Period, short  $<3.5$  hrs  
Orbital evolution, peculiar, spin-orbit  
coupling  
Known Black Widows, 40



MSP  
BINARIES

## Redbacks

“Redback” pulsars,

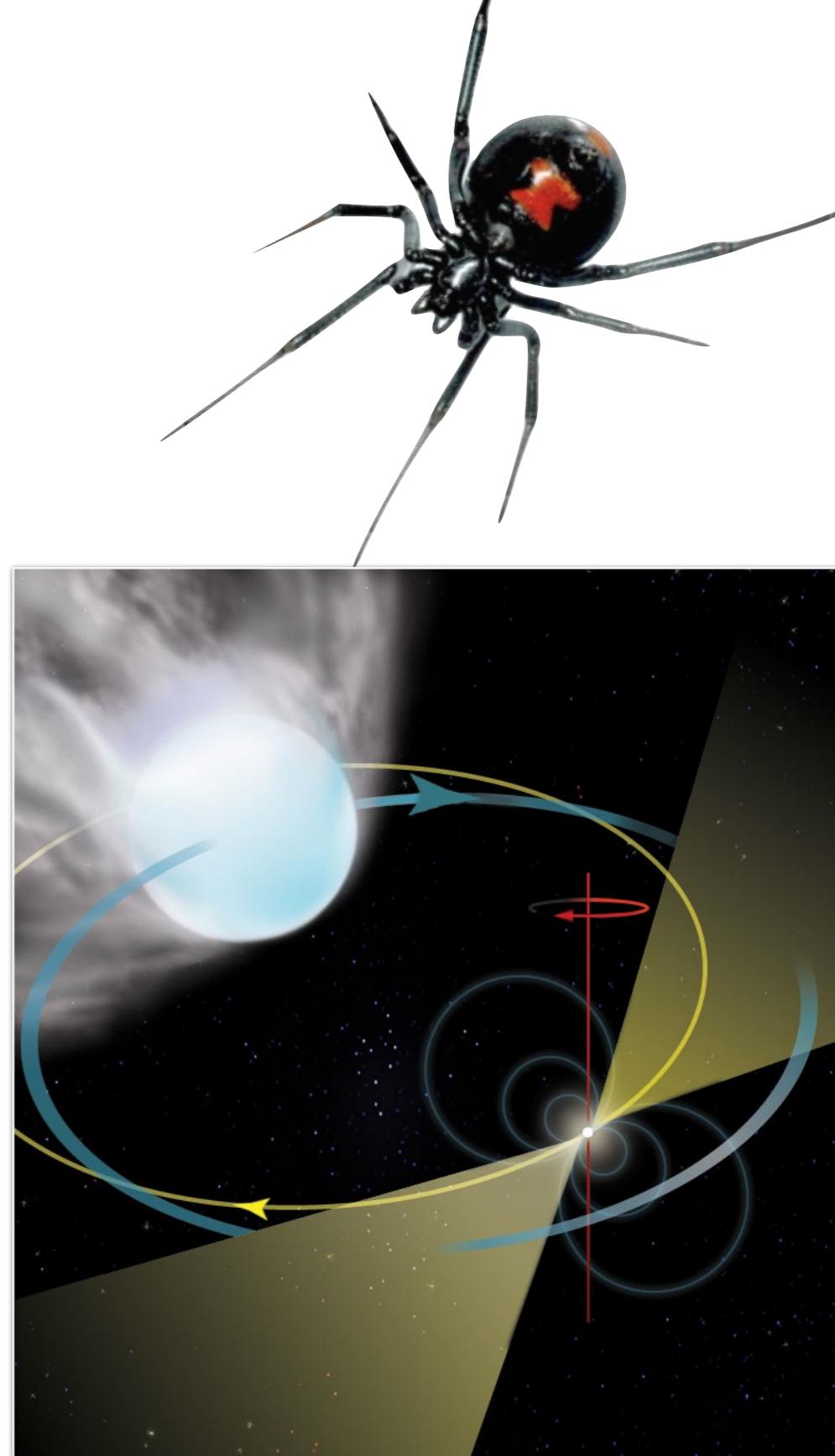
Companion Mass,  $0.15\text{-}0.4 M_{\odot}$

Orbital periods, 4–15 hr

Non-deterministic orbital  
variations

Almost Roche Lobe Filling, Breton  
et al.(2010)

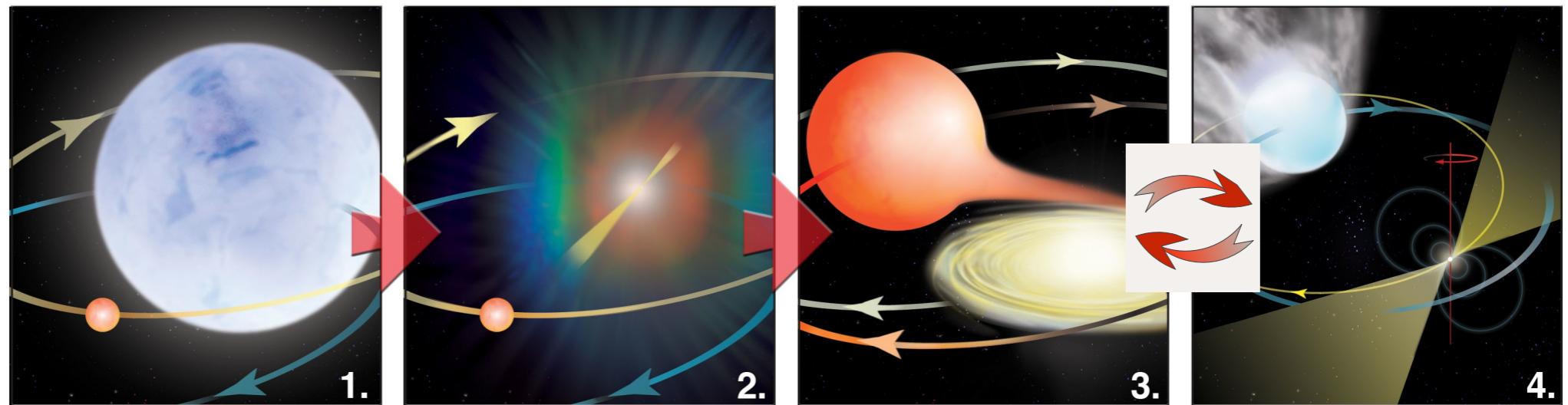
Known Redbacks, 22



## MSP BINARIES

# Transitional Millisecond Pulsars

Bill Saxton; NRAO/AUI/NSF



Binary

SN

Low Mass X-ray  
Binary State

Radio Millisecond  
Pulsar State

**PSR J1023+0038, (1.69 ms)** Archibald et al. (2009)

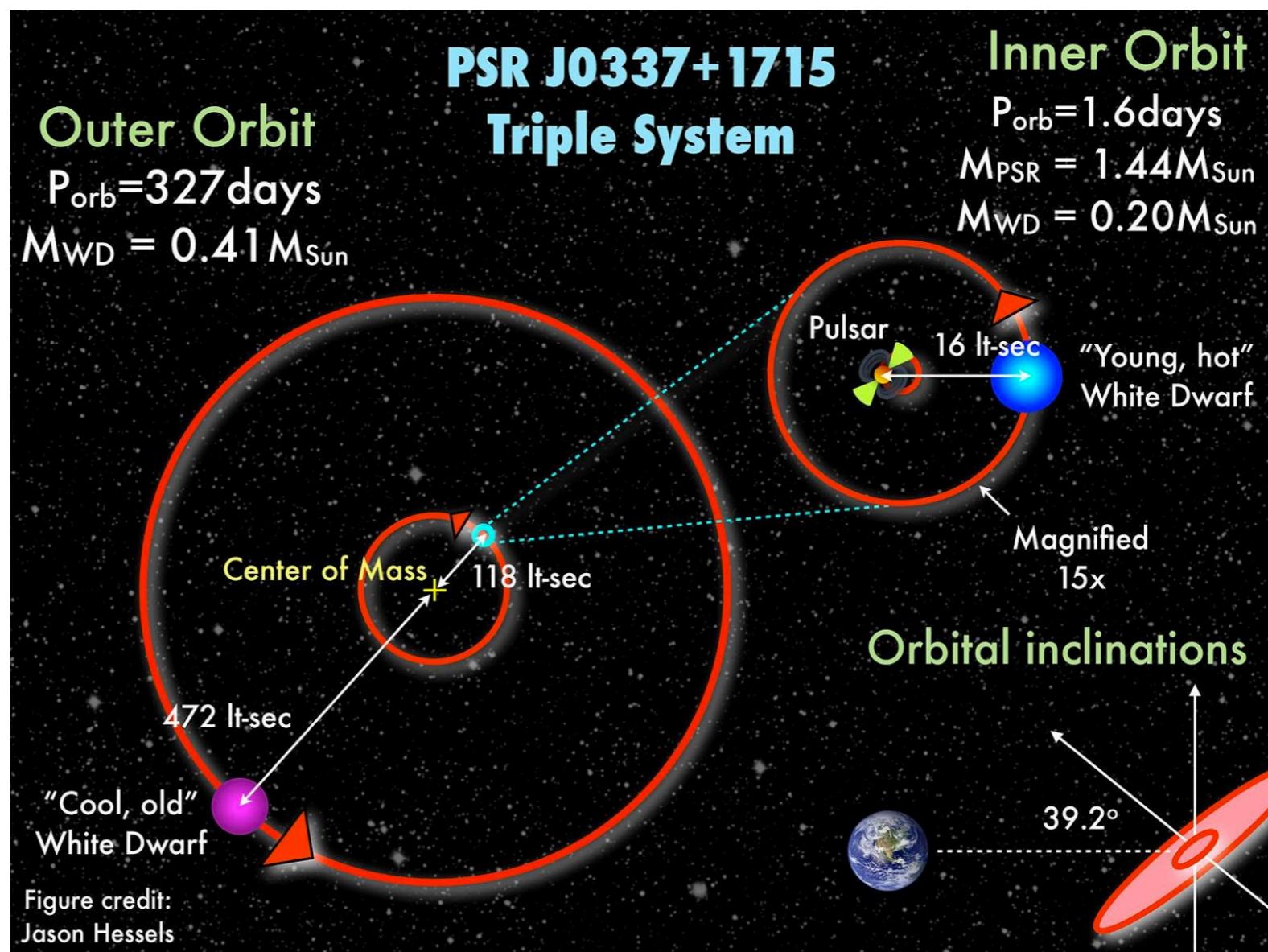
**XSS J12270–4859, (1.69 ms)** Bassa et al. (2014)

**PSR J1824–2452I (3.9 ms)** Papitto et al. (2013)

**3FGL J1544-1125 (?? ms)** Bogdanov & Halpern (2015)

## MSP BINARIES

# Triple MSP System

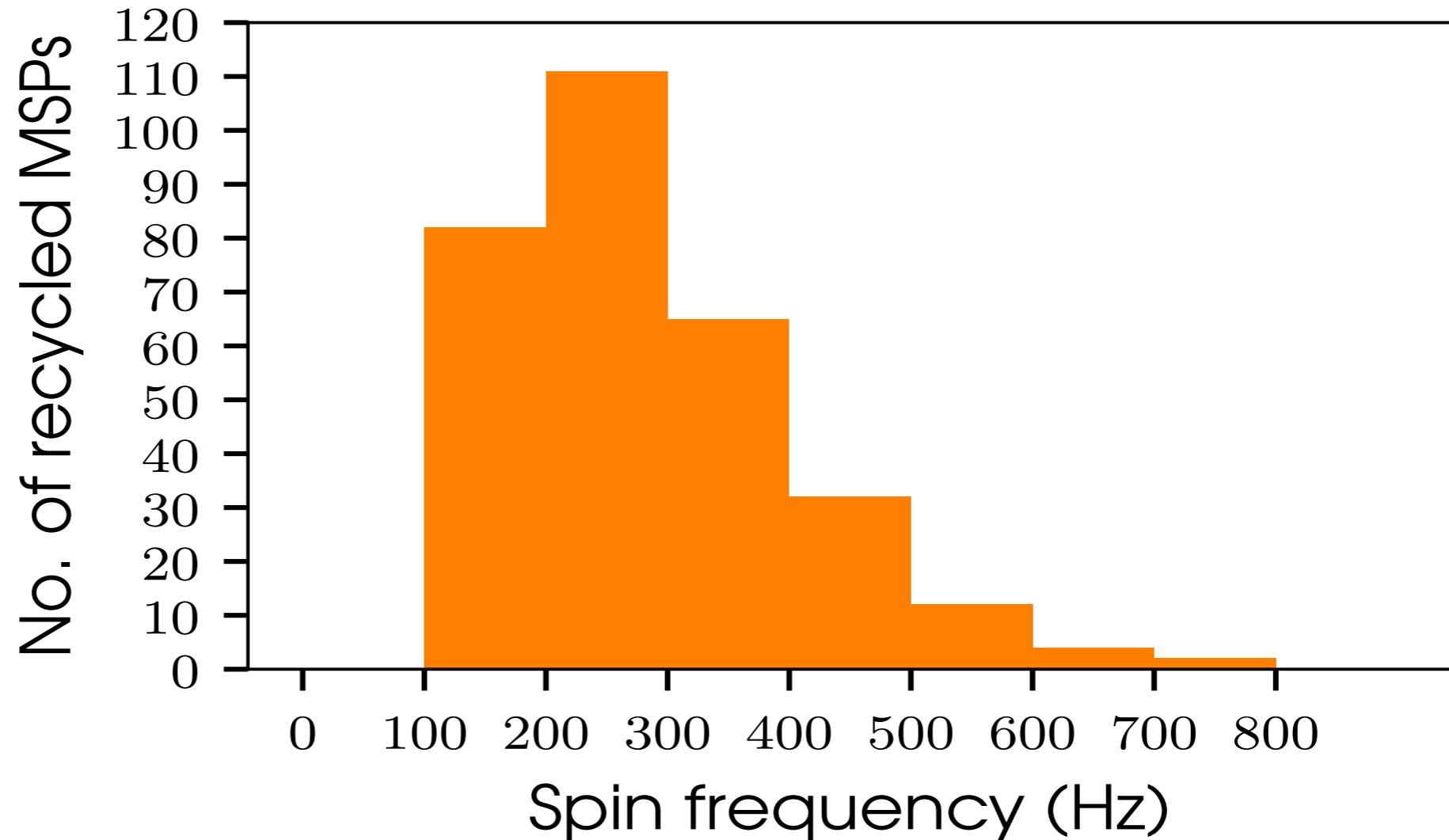


**PSR J0337+1715: 365.95 Hz**

See talk by Anne Archibald

OUTLIERS

# Spin frequency Distribution



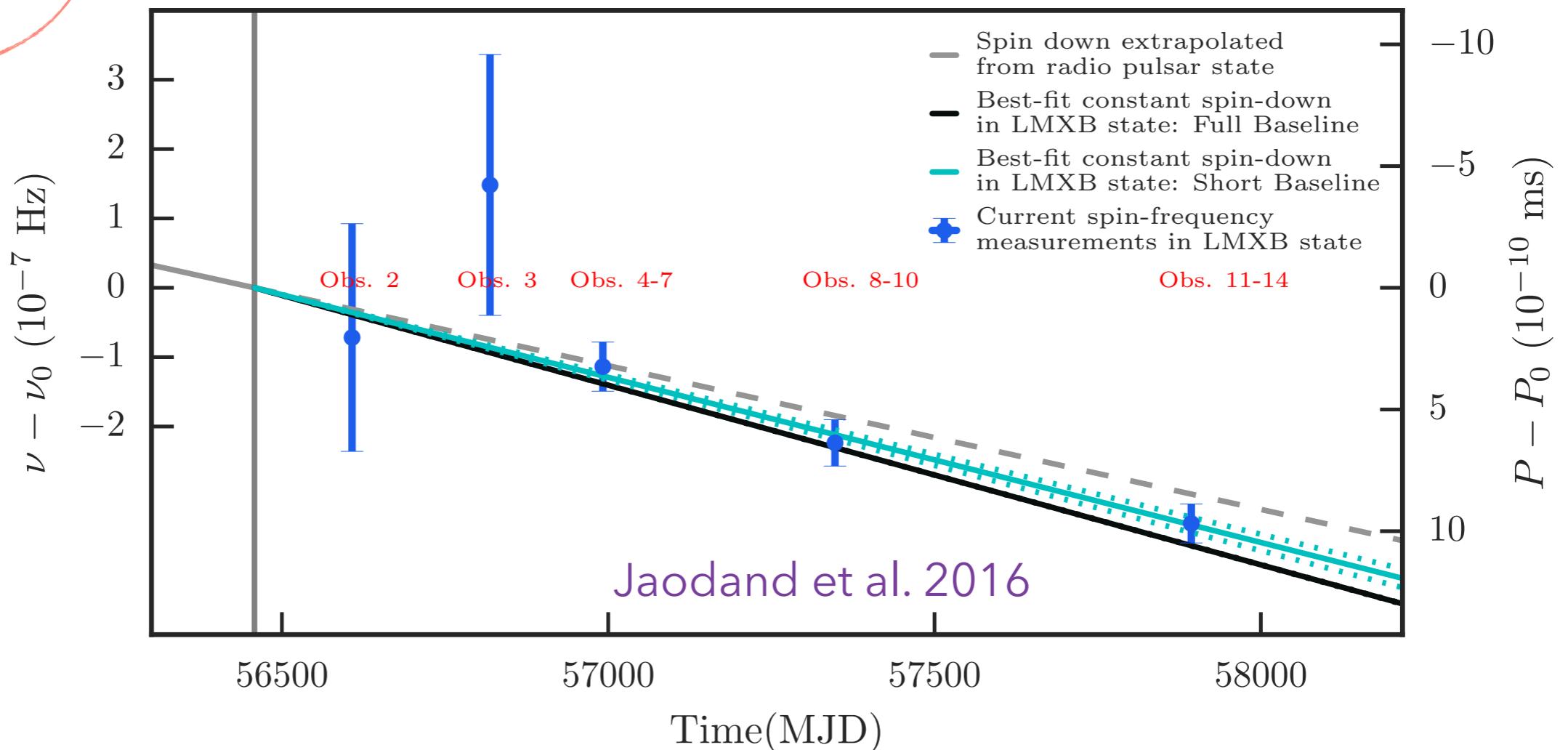
Observational bias or formation mechanism?

Bassa et al 2017 - 707 Hz first galactic field MSP

Plenius et al. 2017 - 411.89 Hz

OUTLIERS

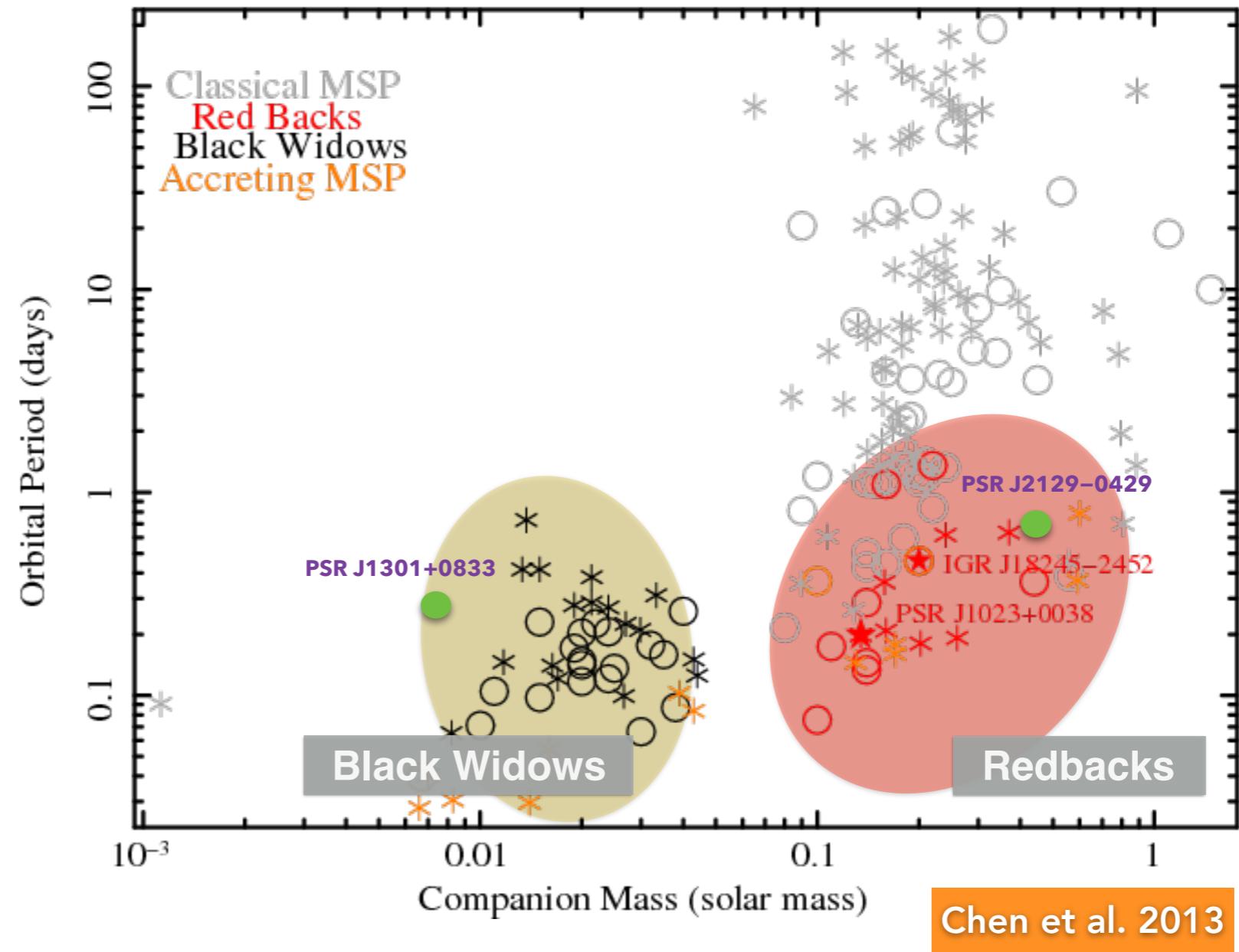
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**OUTLIERS**



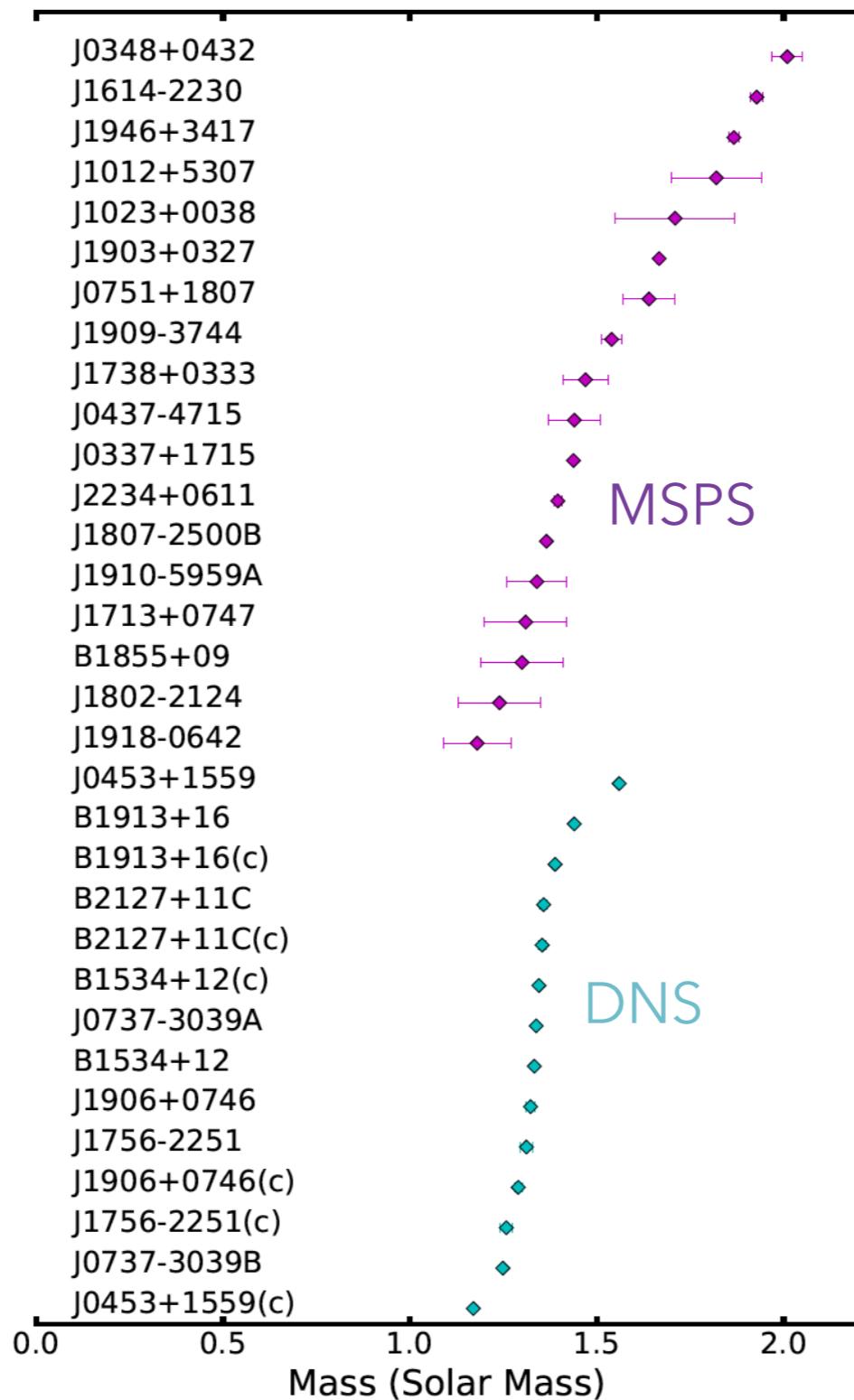
Extreme Black Widows – **Tidarreans: PSR J1301+0833**, Companion Mass  $< 0.01 M_{\odot}$

Romani et al. 2016

Redback at bifurcation – **PSR J2129-0429**, Companion Mass  $= 0.44 M_{\odot}$ ,  $P_{\text{orb}}=15.2 \text{ hr}$

Bellm et al. 2015

# OUTLIERS



PSR J0348+0432 :  
2.01 +/- 0.04 ,M\_sun

Bimodal mass distribution:  
With low mass ( $1.39 M_{\text{sun}}$ )  
and high mass ( $1.8 M_{\text{sun}}$ )

- 1) Precise known mass
- 2) Known total mass
- 3) Known mass ratio

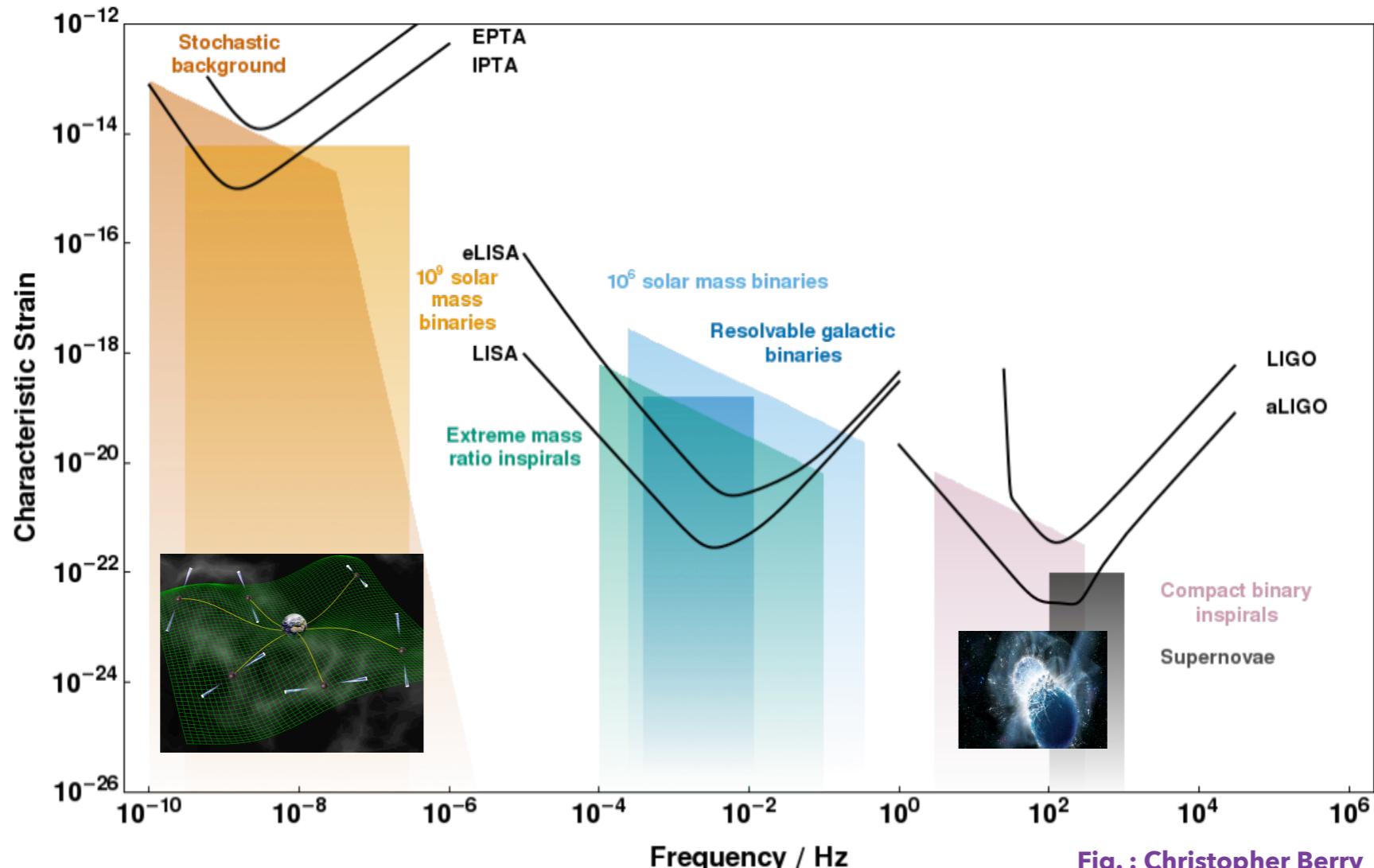
Antoniadis et al. 2013, Antoniadis et al. 2016

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# Applications

Precise Timers & Strongly Self-Gravitating

# PULSAR TIMING ARRAYS



LIGO-VIRGO: gravitational waves (GWs) from BBHs and BNSs (GW170817)

Fig. : Christopher Berry

**PTAs probe GWs in the nHz regime from:**

- Cosmic strings Sanidas et al. 2012, Jose J. Blanco-Pillado et al. 2017
- Supermassive black hole merger - stochastic signal and merger event

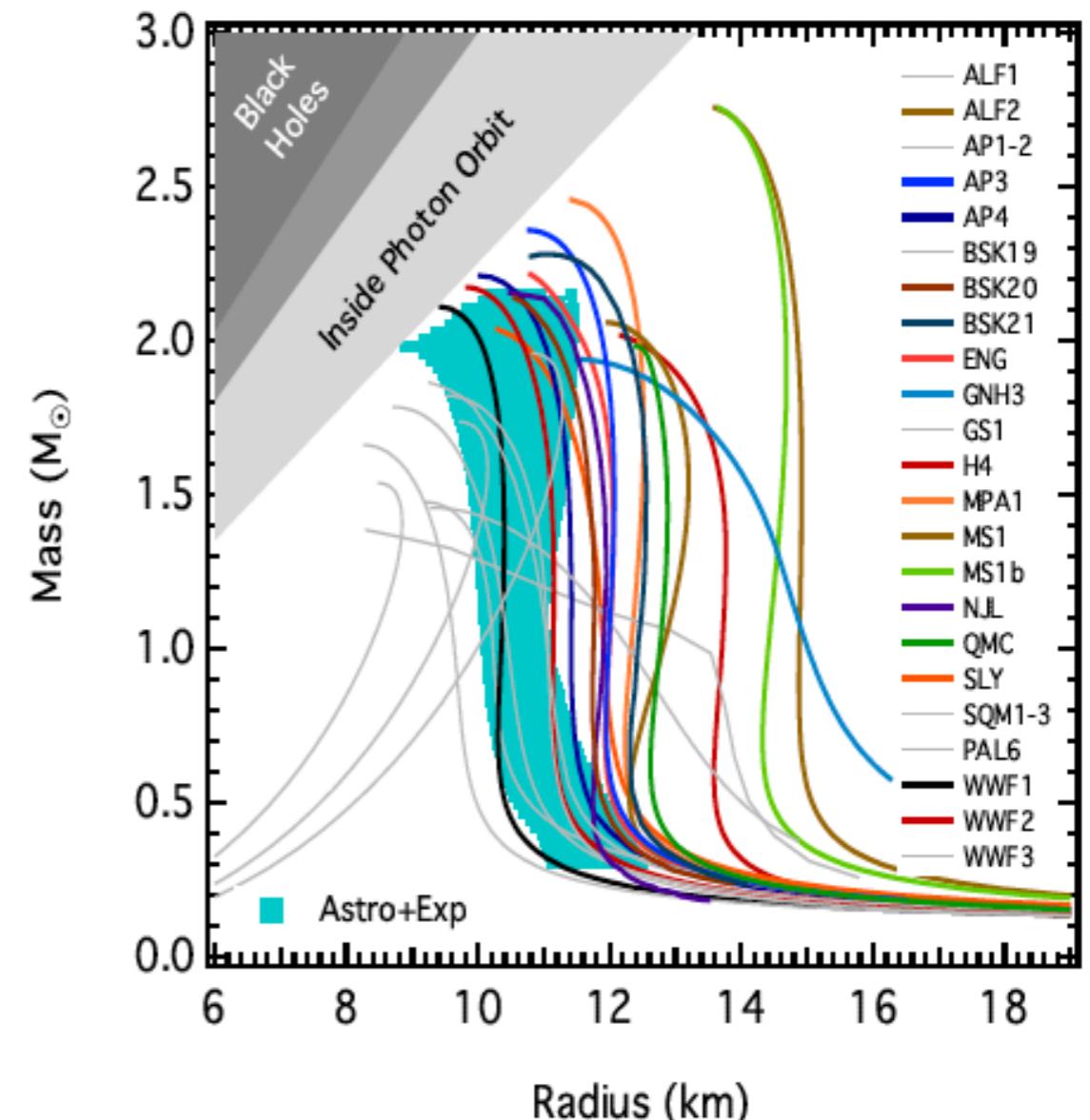
Lentati et al. 2015, Verbiest et al. 2016, Mingarelli et al. 2017

See talk by H. Middleton

# EQUATION OF STATE

Millisecond pulsar mass measurement:  
Fantastic timers, low orbital eccentricities

1. Pulsars in GCs - eccentricity due to interaction  
**Ransom et al. 2005, Freire et al 2008**
2. Shapiro delay measurement  
**Jacoby et al. 2005, Demorest et al. 2010**
3. Spectroscopic measurement - Balmer lines  
**Callanan et al. 1998, Antoniadis et al. 2012**
4. Stellar triples  
**Champion et al 2008, Ransom et al. 2014**



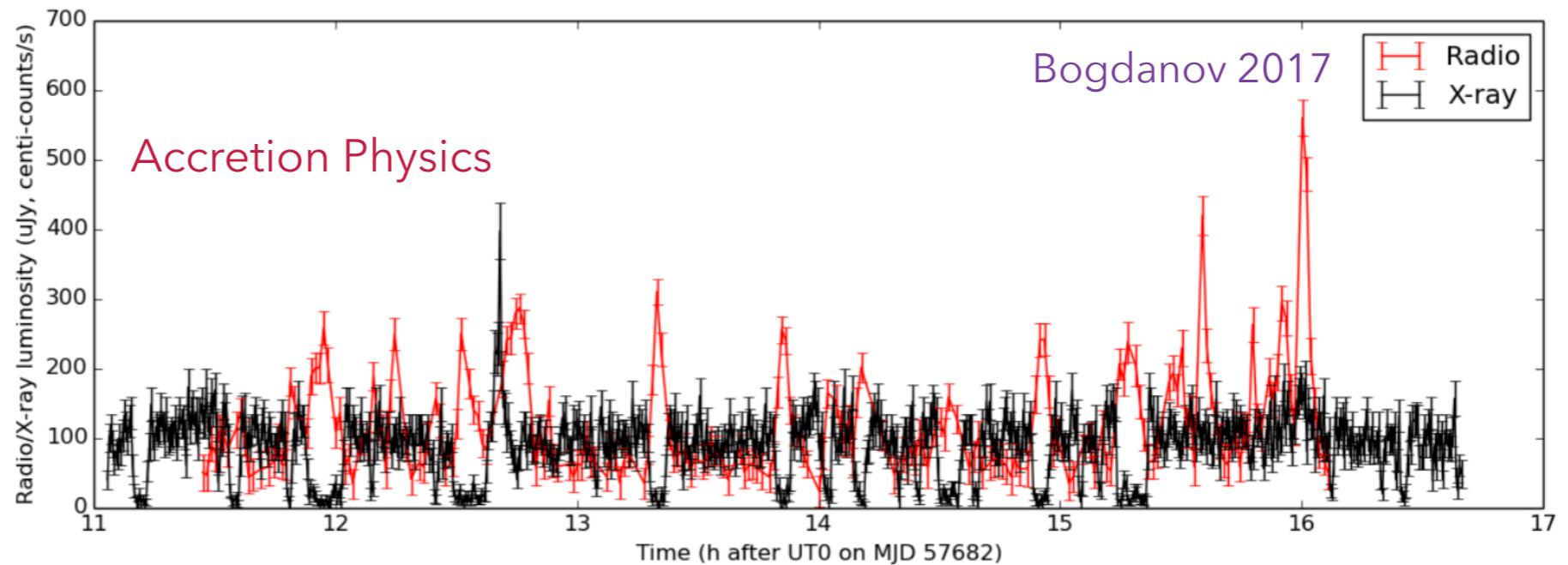
Millisecond pulsar radius measurement:

1. Spectroscopic measurement
2. Timing Measurements

**Review: Ozel and Freire 2016**

See talk by N.U.Bastian for more details

## OTHER APPLICATIONS



**tMSPs are fantastic probes of low-level accretion onto neutron stars**

Jaodand, Hessels & Archibald 2017

Accreting millisecond X-ray pulsars

Review: Patruno and Watts 2012

Understanding of pulsar emission physics  
Galactic centre excess?

**See talks by:**

Stellar Evolution - John Antoniadis

Inter stellar medium - Caterina Tiburzi

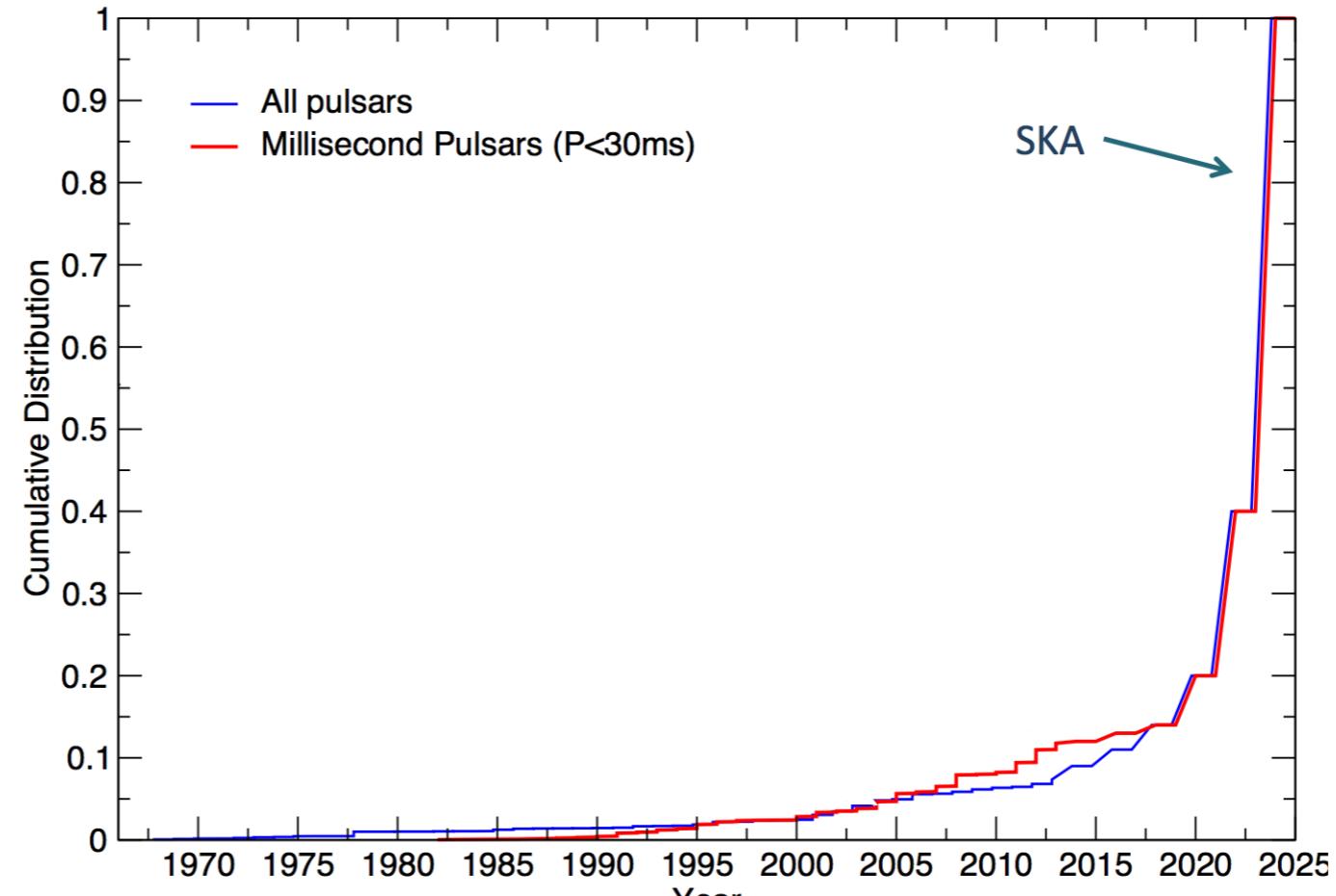
Strong field tests of gravity - Shao and Archibald

4

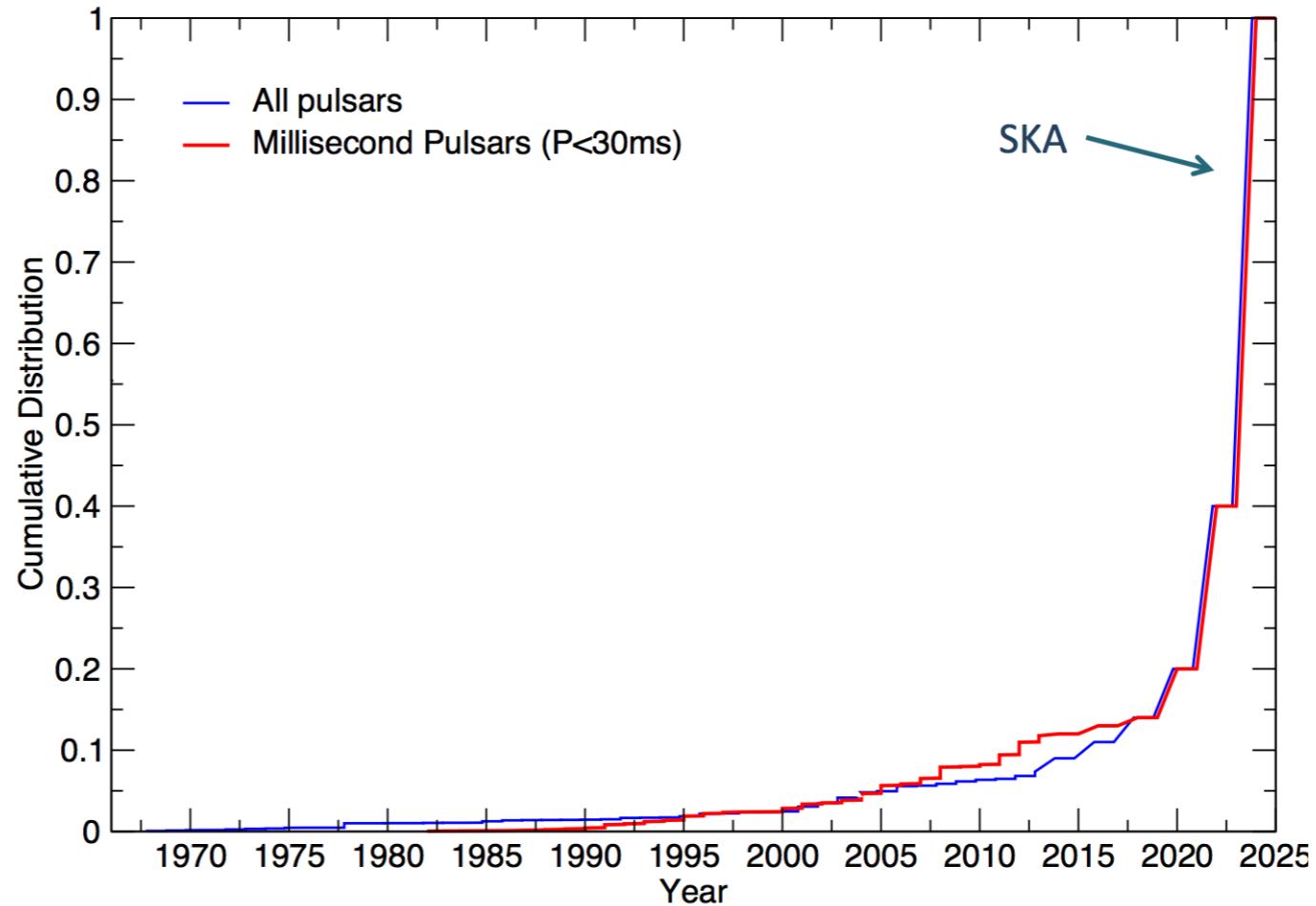
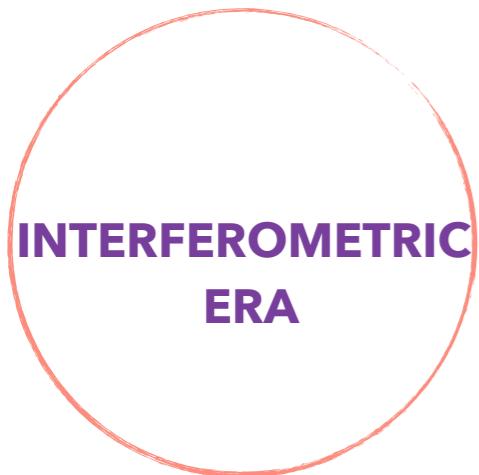
## Future Prospects

New Means To Find MSPS & Holy Grails

**INTERFEROMETRIC  
ERA**

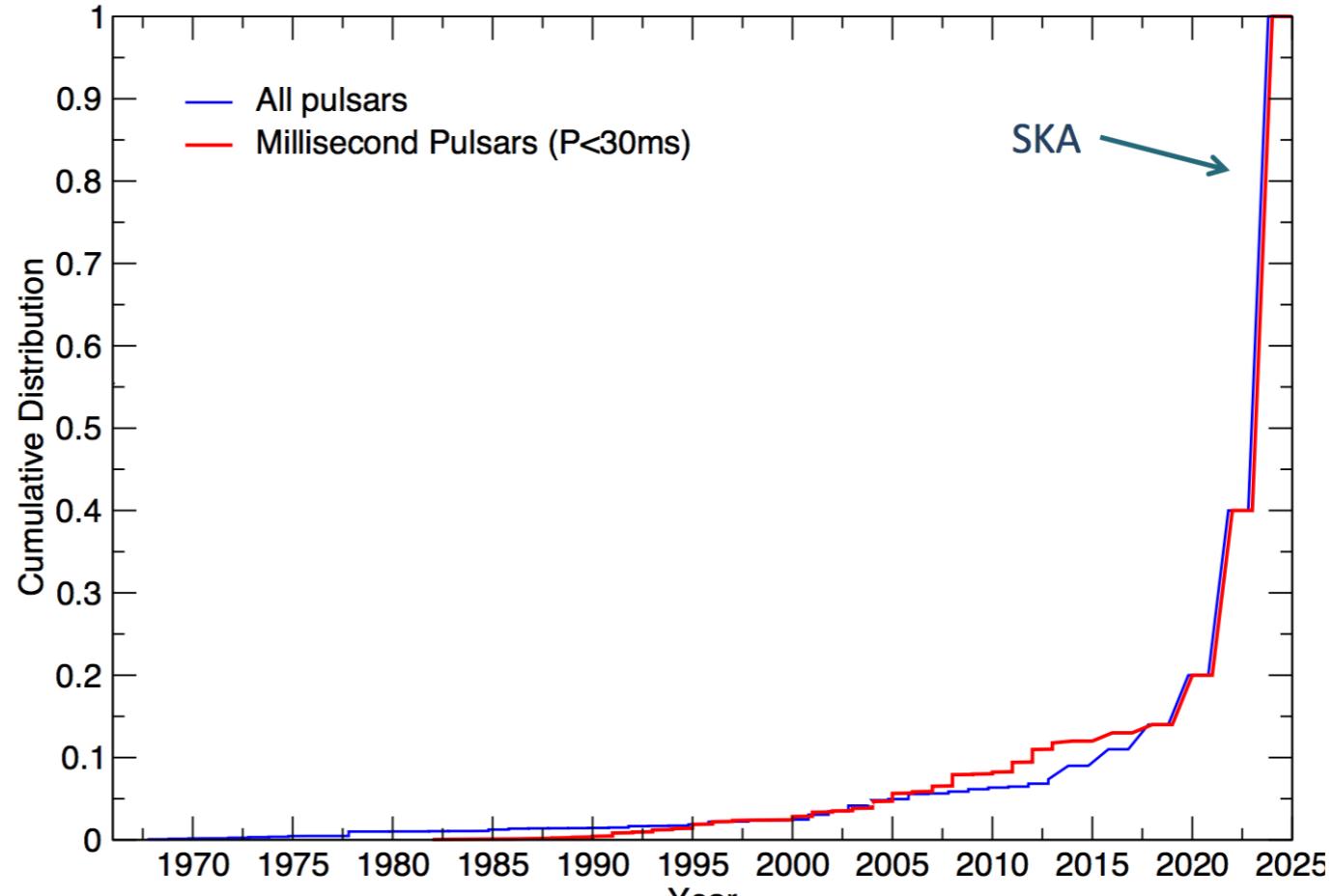
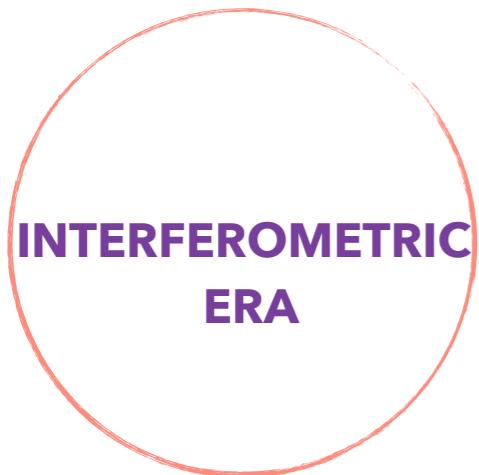


**Fig: M. Kramer**



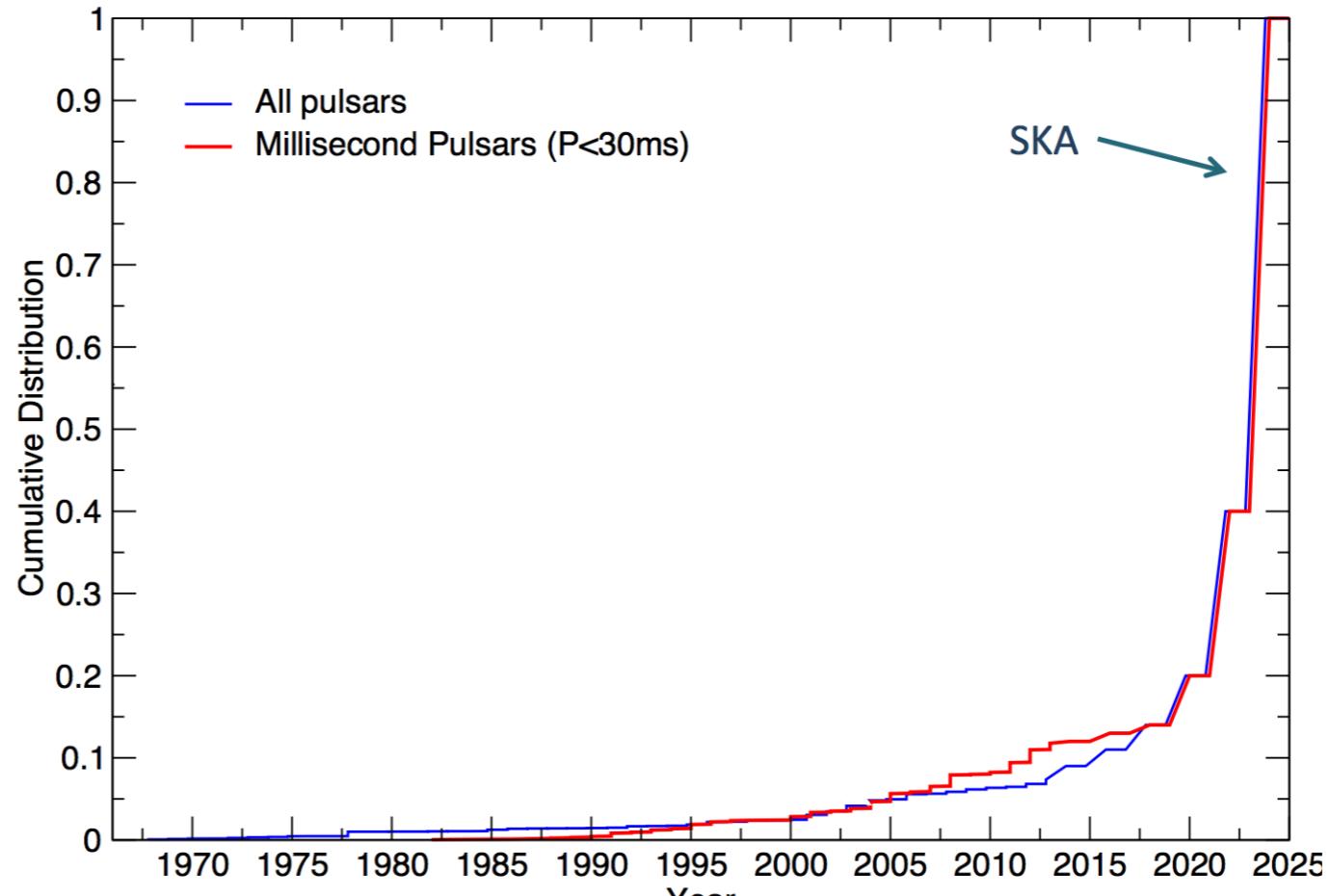
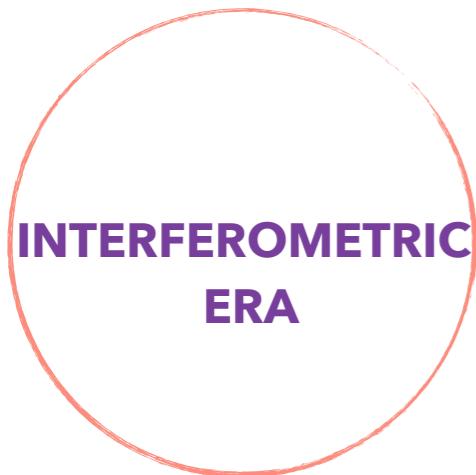
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Higher instantaneous sensitivities for SKA, MeerKAT and pathfinders



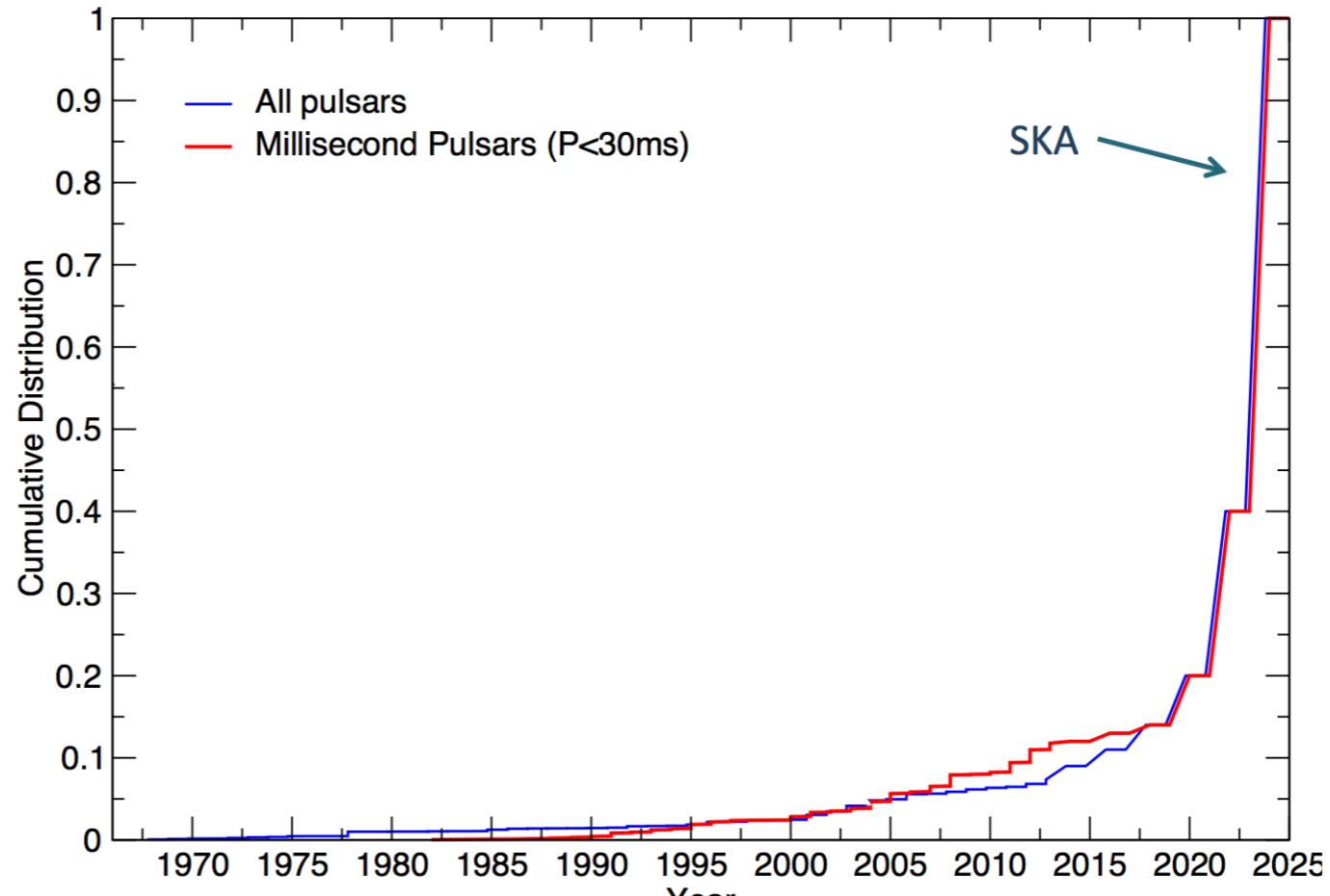
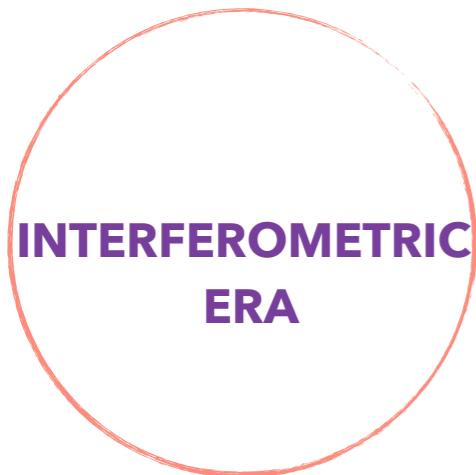
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SKA's large field of view, collecting area, bandwidth



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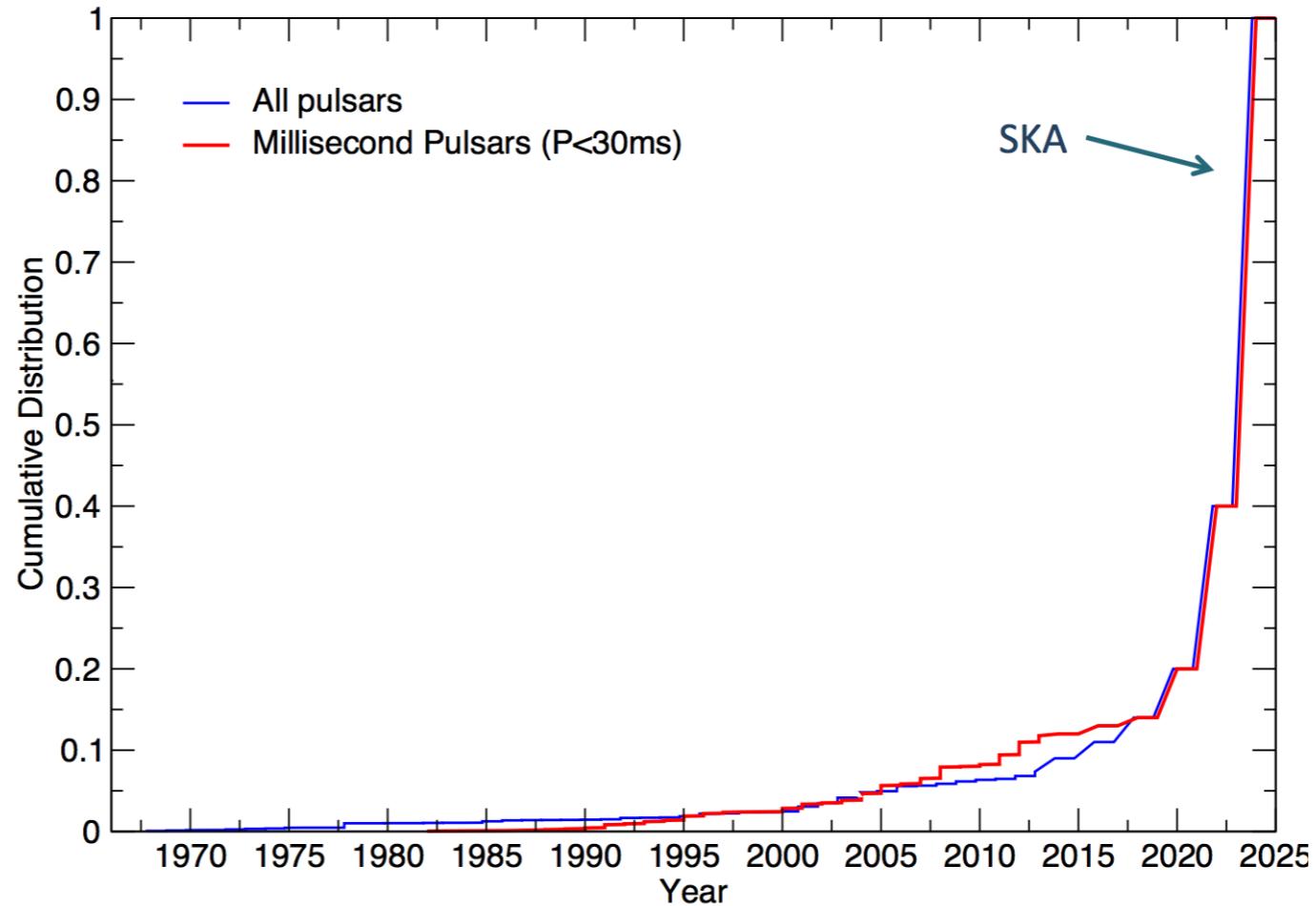
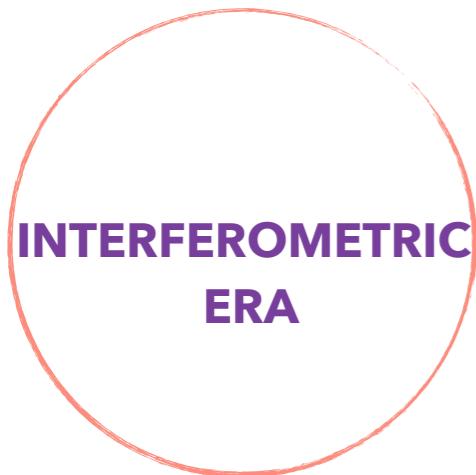
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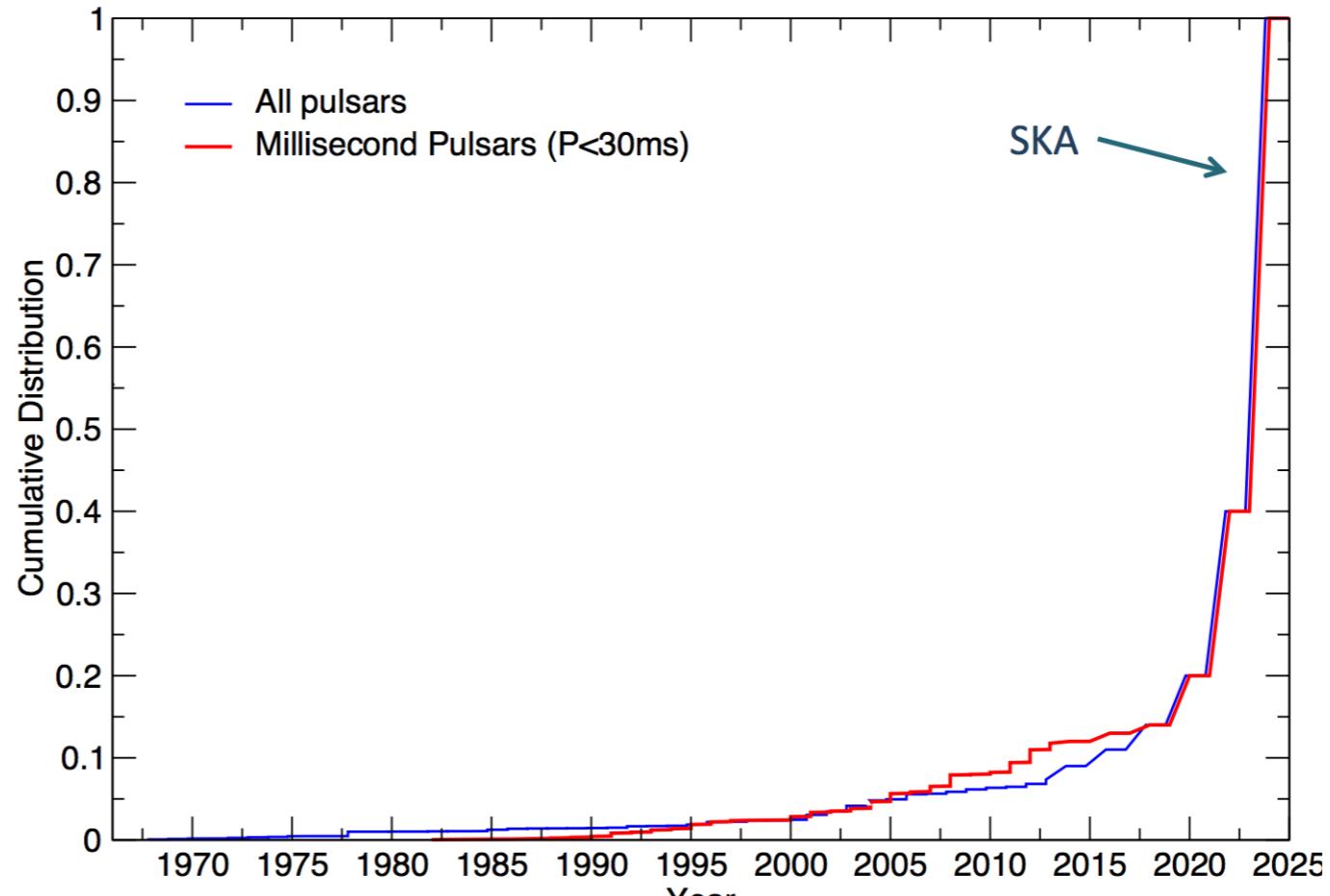
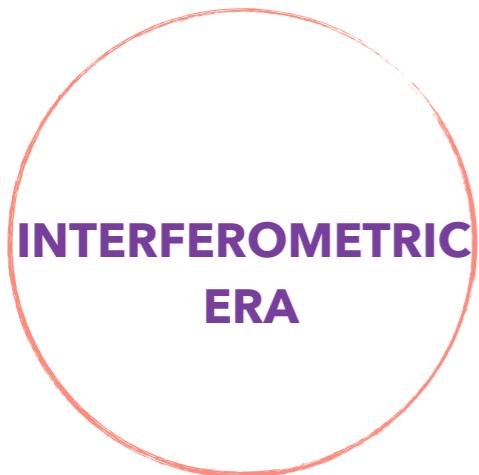
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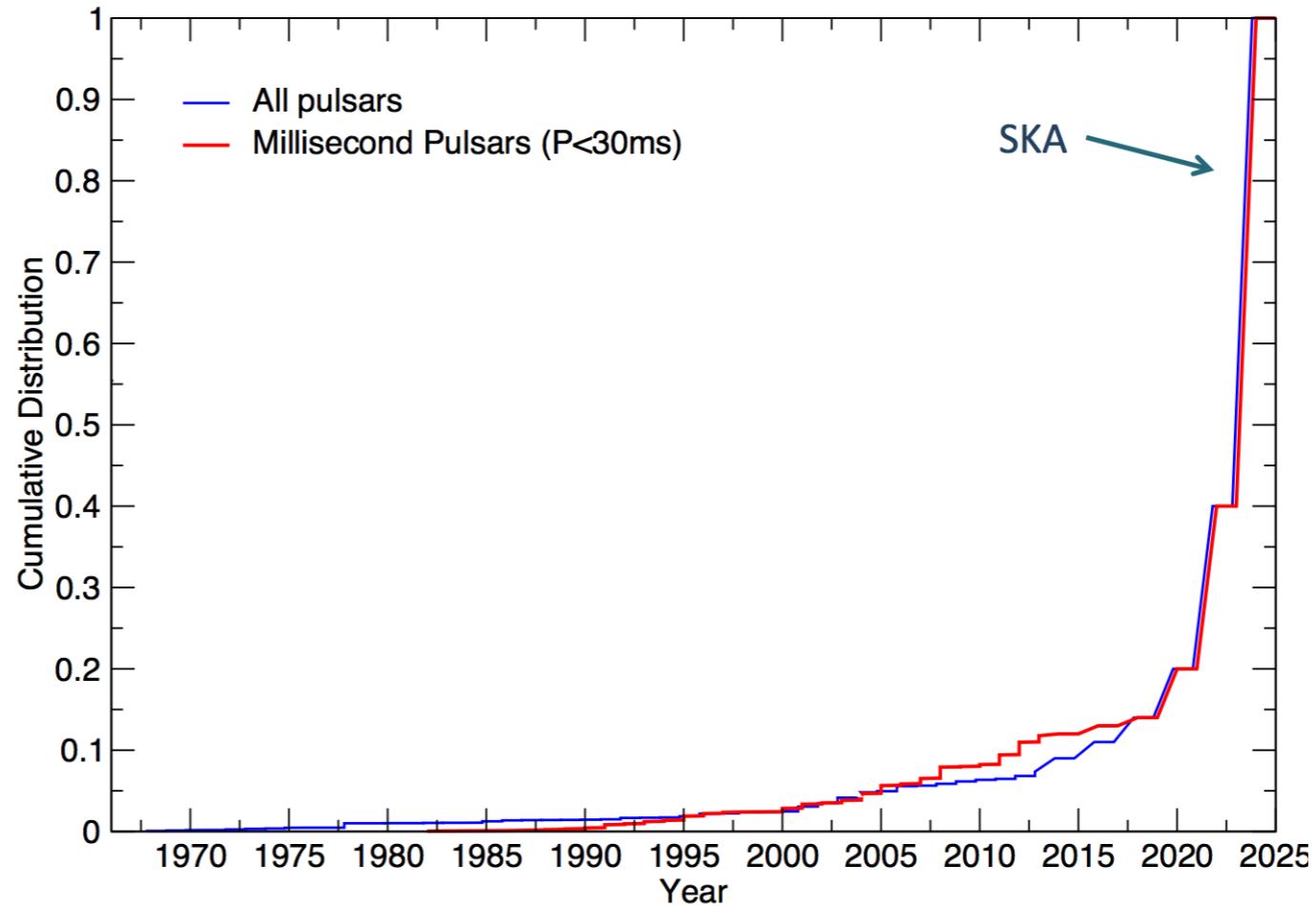
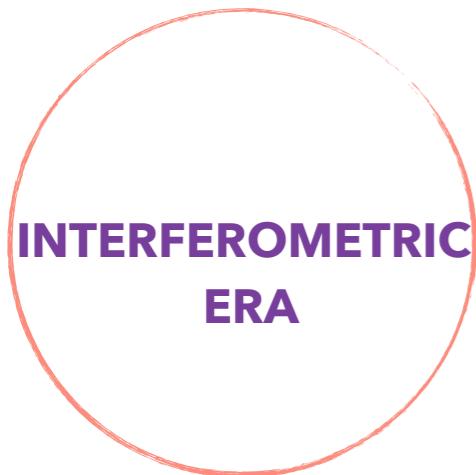


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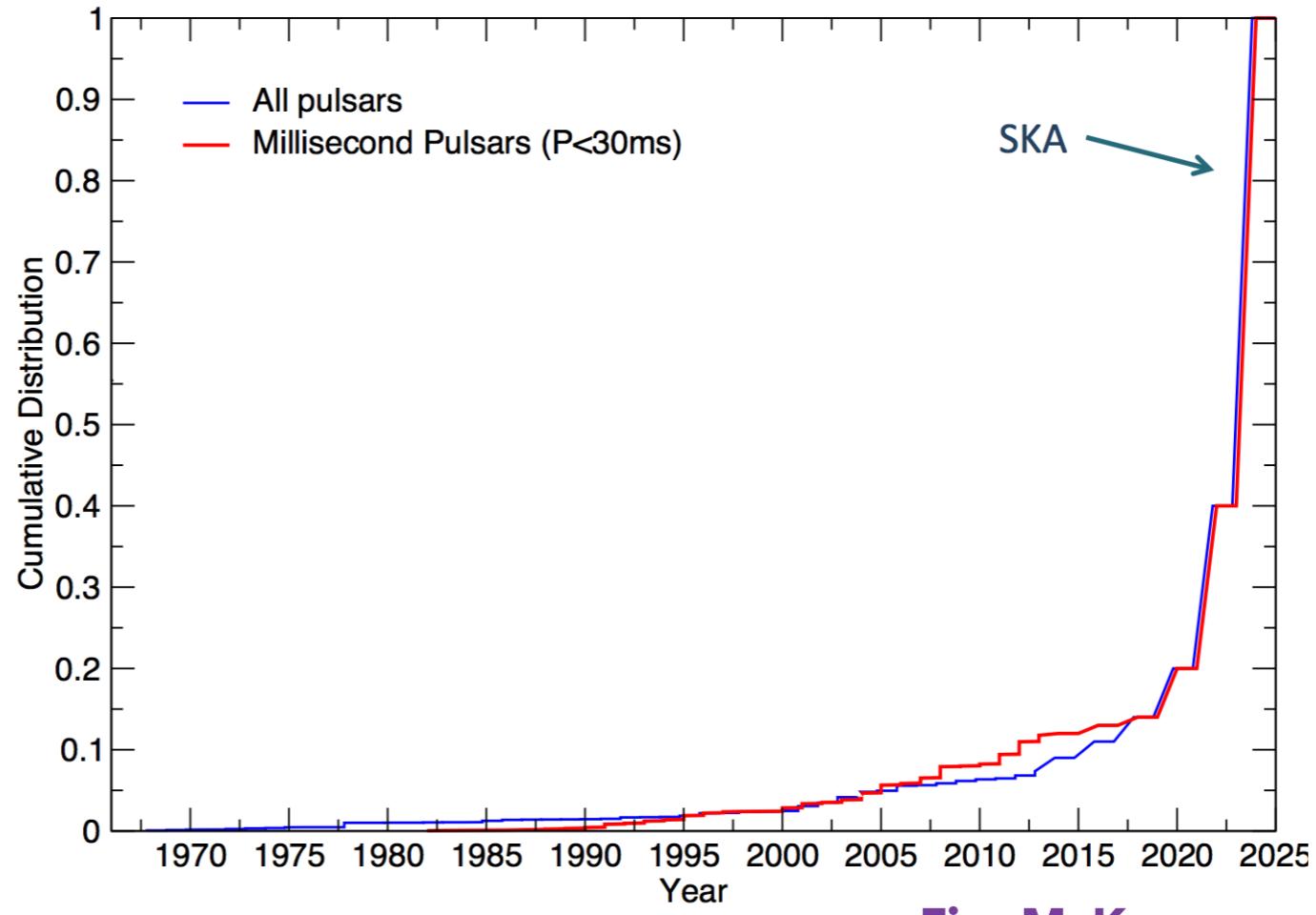
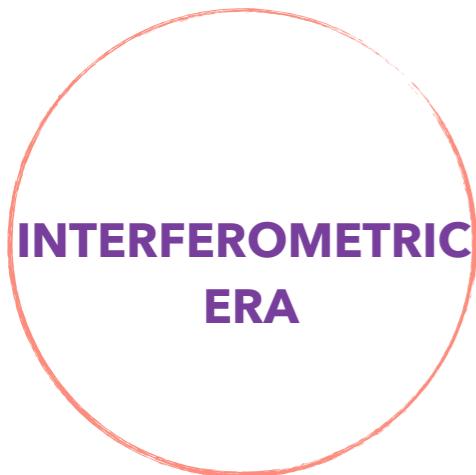


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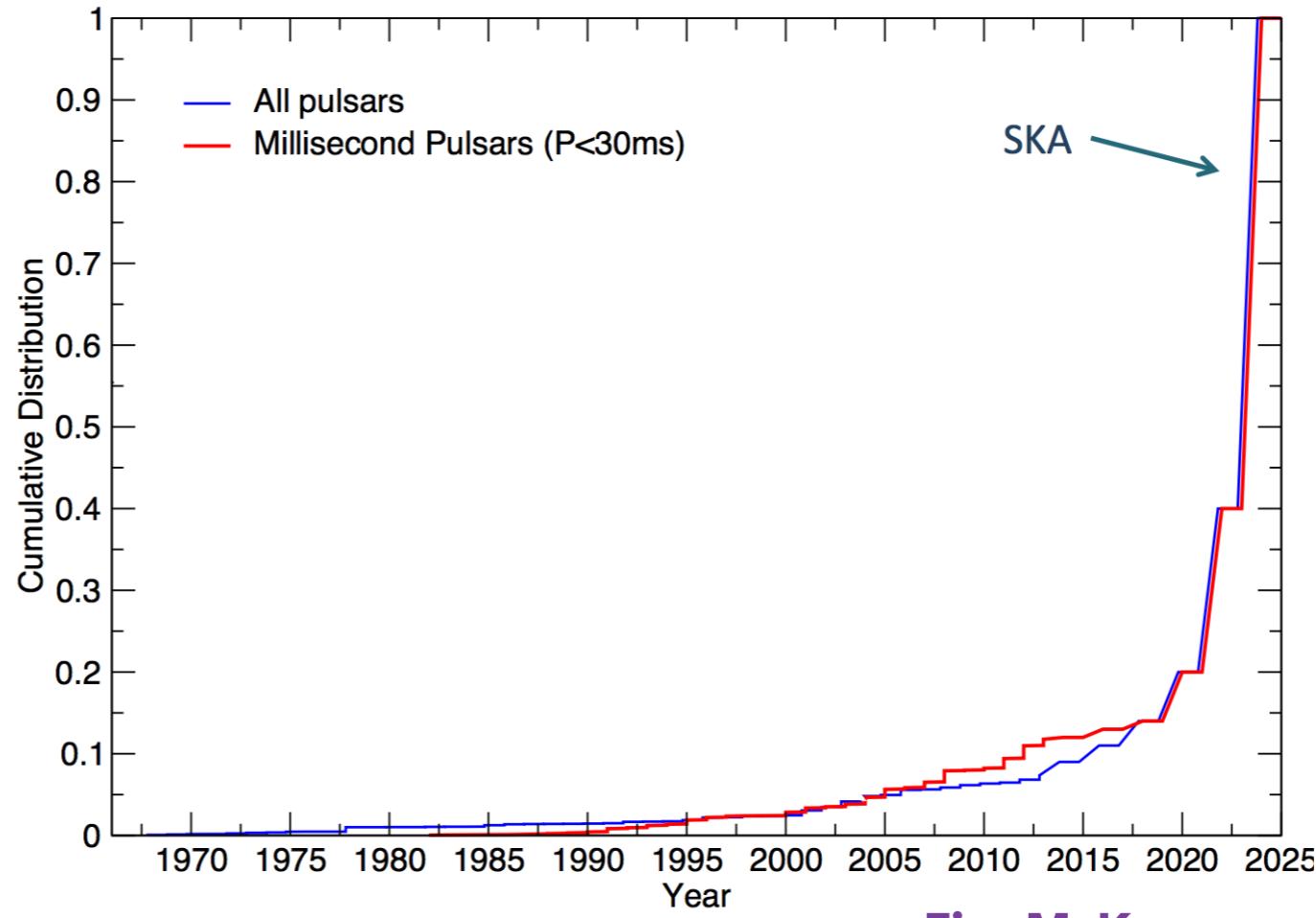
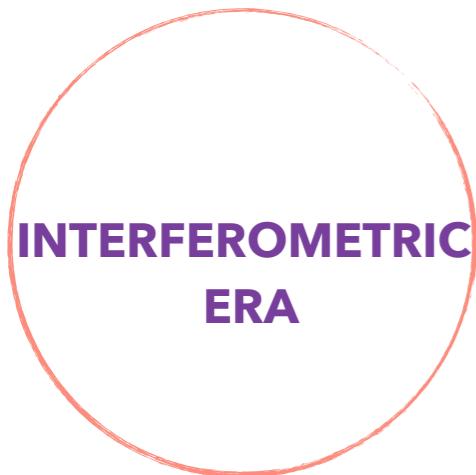
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Astron  
Hackathon

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**NEW SKY  
SURVEYS**



## NEW SKY SURVEYS

- **VLA Sky Survey (VLASS)** - 3 epochs of high cadence observation (4504 hr, 34000 deg<sup>2</sup>, 6 sources at SNR of 10/deg<sup>2</sup>)
  - First epoch commenced in this year in Sep.
  - 2-4GHz survey : Flat as well as modest spectrum sources

[VLASS white papers , Bhakta et al. 2017](#)



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  - **NICER**

[Gendreau and Arzoumanian, Nature Astronomy 1, \(2017\)](#)



**HOLY GRAILS**

## Fantastic pulsars and how to find them?

Formation Mechanisms? Specific observation strategies.



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Sub-ms pulsars

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Ultracompact binaries?  $P_{\text{orb}} < 1\text{hr}$

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MSP with planets

Formation Mechanisms? Specific observation strategies.

## SUMMARY

- Population of MSPs has nearly doubled since 2010
- We expect a new burst in MSP discoveries with new surveys such as SKA
- New discoveries lead to diversity and **diversity in pulsar population translates to fantastic new probes of physics!**
- Extragalactic pulsars?
- Creative use of MSPs - scintillometry -
- Open to new finds during MSP searches such as FRBs and unknowns

# Acknowledgments

- Plots supported by Matt Pitkin's Psrqpy package
- ATNF pulsar catalogue
- P. Freire's globular cluster catalogue
- D. Lorimer's Galactic MSP catalogue
- A.Patruno's MSP catalogue
- *Jason Hessels for guidance!*
- **Wonderful reviews**