German ALMA Community Days 2015

(Bonn, 25-27 March 2015) Toma Bădescu – Argelander Institut für Astronomie



EUROPEAN ARC ALMA Regional Centre || Germany

Brief outline:

- What are we measuring (brief theoretical discussion)
- How to use the sensitivity calculator
- What to look out for
- Minor wrap-up





$$P_{
u} = I_{
u} dAd\Omega$$

 $= \frac{2kT_{src}}{\lambda^2} dAd\Omega$ RJ limit means: power ~ specific intensity
~ (brightness) temperature
 $T_{sys} = T_{sky} + T_{Rx}$ Something we don't need!

$$T_{sys} > > T_{src}$$

We need $T_{src}!$ What to do?

$$T_{sys} >> T_{src}$$
 This is usually the case!
 $T_{src} > T_{rms}$ This is what we need to get a measurement!
 $T_{rms} = \frac{T_{sys}}{\sqrt{N}}$ Reduce T_{rms} by sampling T_{sys} ...

$$T_{rms} = rac{T_{sys}}{\sqrt{N}}$$
 The noise when measuring T_{sys} decreases with increasing N (sampling points)

But what is N?

N ~ n(n-1)/2 = number of baselines, n is the number of antennas

- $\sim \Delta v$ = frequency range (bandwidth) in Hz
- ~ t = exposure time (on source time) in seconds
- ~ 2 = a factor for double polarization mode

$$T_{rms} \simeq \frac{T_{sys}}{\sqrt{n(n-1)} \cdot \Delta \nu \cdot t} [K]$$

Back to flux and Jansky... or milliJansky (mJy):

$$\text{SEFD} = \frac{T_{sys}}{(K/Jy)} = \frac{T_{sys}}{A_e/2k}$$

System Equivalent Flux Density

$$S_{rms} \simeq \frac{SEFD}{\sqrt{n(n-1)} \cdot \Delta \nu \cdot t} [Jy]$$

 $n(n-1) \sim number of baselines, n is the number of antennas$ $<math>\Delta v =$ frequency range (bandwidth) in Hz t = exposure time (on source time) in seconds

Inputs to calculator:

- declination (for elevation)
- frequency (e.g. for atmospheric absorbtion etc.)
- weather (optional)
- number of antennas (gives number of baselines)
- resolution (gives the Δv)
- integration time (gives the value of t)

Output:

- Sensitivity reached



Inputs to calculator:

- declination (for elevation)
- frequency (e.g. for atmospheric absorbtion etc.)
- weather (optional)
- number of antennas (gives number of baselines)
- resolution (gives the Δv)
- desired sensitivity (gives the value of S_{rms} or T_{rms})

Output:

- Integration time required



What to look out for?

Atmospheric absorbtion:



Source size:



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Notes on calibration:



Thank you, and make sure to check the summary slides in the afternoon

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