

## OBSERVATIONAL COSMOLOGY

### PROBLEM SHEET 2 - 04/05/2017

Supernovae Ia are used to derive the Hubble diagram and provide strong evidence for the acceleration of cosmic expansion. The key fact is that they are *standardizable candles*. In fact, observing their light curves (i.e. the evolution of light intensity as a function of time) an empirical relationship was found between the maximum luminosity and the width of the curve itself. It can be expressed as

$$M_{max} = a \cdot m_{15} + b,$$

where  $M_{max}$  is the peak magnitude and  $m_{15}$  is the decline of the magnitude 15 days after it has reached its peak.

1) We want to design a new experiment to measure  $a$  and  $b$  with great accuracy. For simplicity, let us assume that we can compress the supernova data in a set of bins which are spaced by  $\Delta m_{15} = 1$  (using some specific units). We expect statistical errors on  $M_{max}$  of  $\sigma_M = 0.1$ , which are Gaussian distributed and independent. Using the Fisher matrix method, determine the **number of bins** we need to consider to get marginal errors  $\sigma_a = 0.01$  and  $\sigma_b = 0.01$  on the linear fit parameters  $a$  and  $b$ .

Hints:

$$\sum_{n=1}^k n = \frac{k(k+1)}{2}$$
$$\sum_{n=1}^k n^2 = \frac{k(k+1)(2k+1)}{6}$$

2) Current experiments are less accurate and give  $\sigma_M = 0.7$ . We collected data in five bins, ranging from  $m_{15} = 1$  to 5 and the results are:  $-17.3$ ,  $-15.5$ ,  $-15.6$ ,  $-12.0$ ,  $-10.1$ . Using the MCMC method, plot the 68% credibility interval in parameter space (both in the joint and in the marginalized cases).